

# Neutrino Oscillations with MINOS



Jeff Hartnell

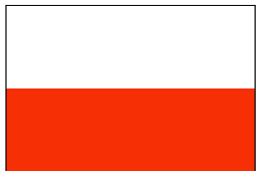
University of Sussex  
for the MINOS collaboration



# Introduction

- MINOS physics goals
- NuMI neutrino beam
- MINOS detectors
- Results:
  - Muon neutrino disappearance analysis
  - Electron neutrino appearance analysis (new!)
  - Neutral current analysis: sterile neutrino mixing
- Future plans





**28 institutions  
140 scientists**

# The MINOS Collaboration

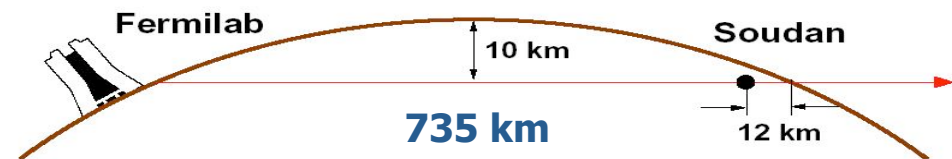


**Argonne • Athens • Benedictine • Brookhaven • Caltech • Cambridge • Campinas • Fermilab  
Harvard • Holy Cross • IIT • Indiana • Minnesota-Twin Cities • Minnesota-Duluth • Otterbein  
Oxford • Pittsburgh • Rutherford • Sao Paulo • South Carolina • Stanford • Sussex • Texas A&M  
Texas-Austin • Tufts • UCL • Warsaw • William & Mary**



# MINOS Overview

- Main Injector Neutrino Oscillation Search
- Neutrinos at the Main Injector (NuMI) beam at Fermilab
- Two detectors:
  - Near detector at Fermilab
    - measure beam composition
    - energy spectrum
  - Far detector in Minnesota
    - search for and study oscillations

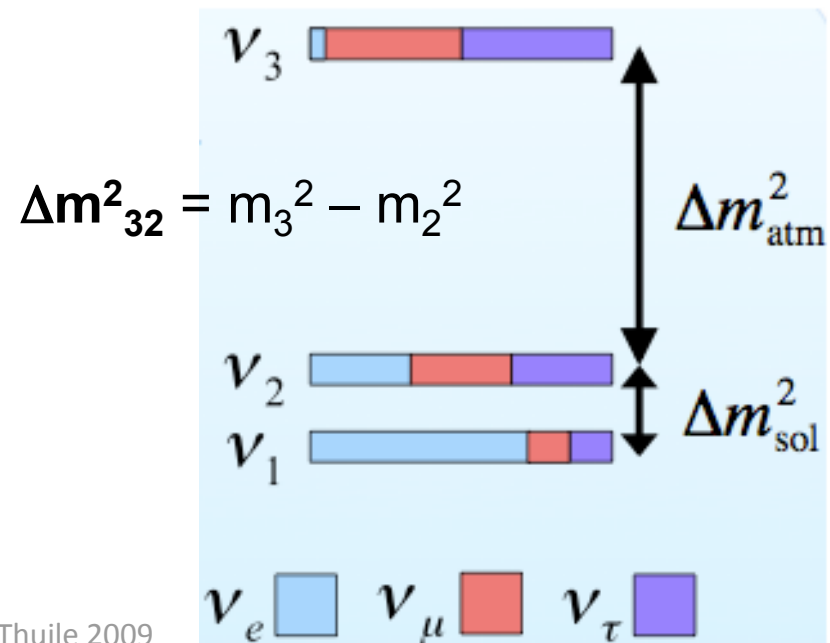


# MINOS Physics Goals

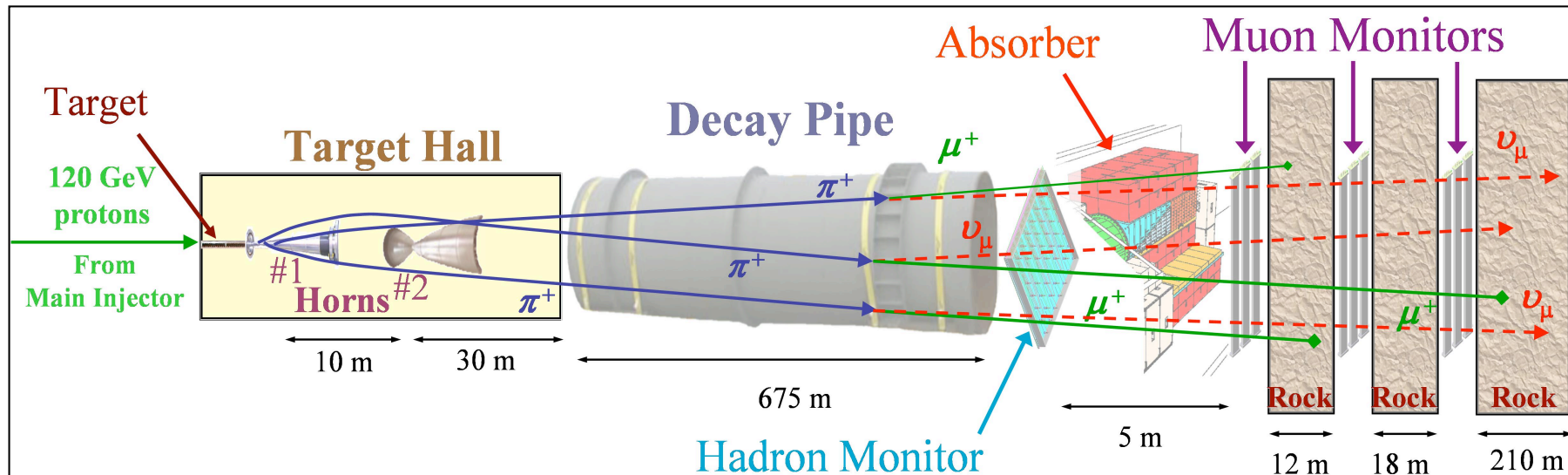
- Test the  $\nu_\mu \rightarrow \nu_x$  oscillation hypothesis
  - Measure precisely  $|\Delta m_{32}^2|$  and  $\sin^2(2\theta_{23})$
- Search for sub-dominant  $\nu_\mu \rightarrow \nu_e$  oscillations
  - sensitive to  $\theta_{13}$
- Other MINOS physics:
  - Search for sterile neutrinos, CPT/Lorentz violation
  - Compare  $\nu_\mu$ ,  $\bar{\nu}_\mu$  oscillations
  - Studies of cosmic rays and atmospheric neutrinos
  - Neutrino interaction studies in the Near detector

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

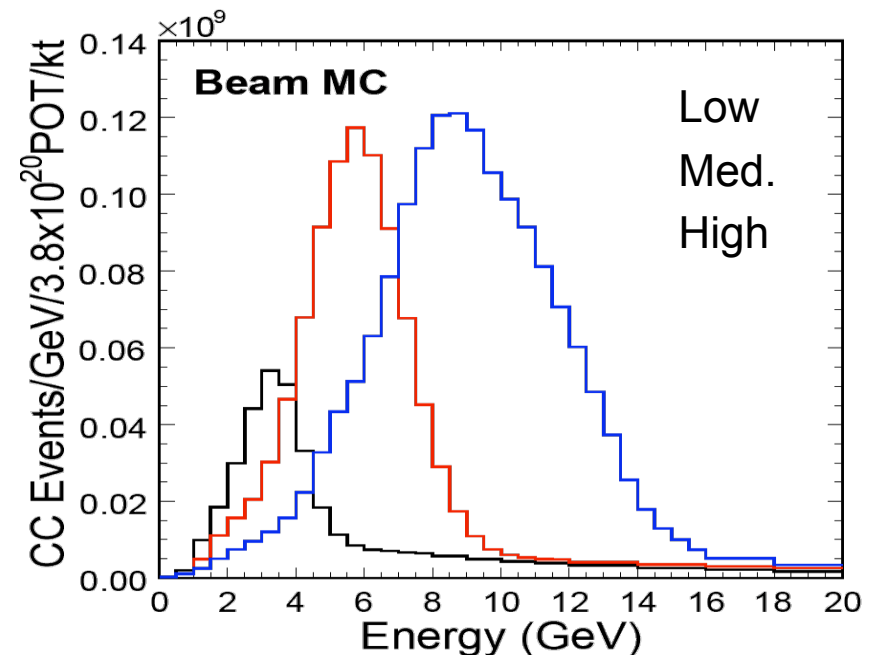
$\nu_e$  appearance  $\downarrow$  (pointing to  $U_{e3}$ )  
 $\uparrow$  (pointing to  $U_{\mu3}$ )  $\nu_\mu$  disappearance



# Neutrino Beam (NuMI)

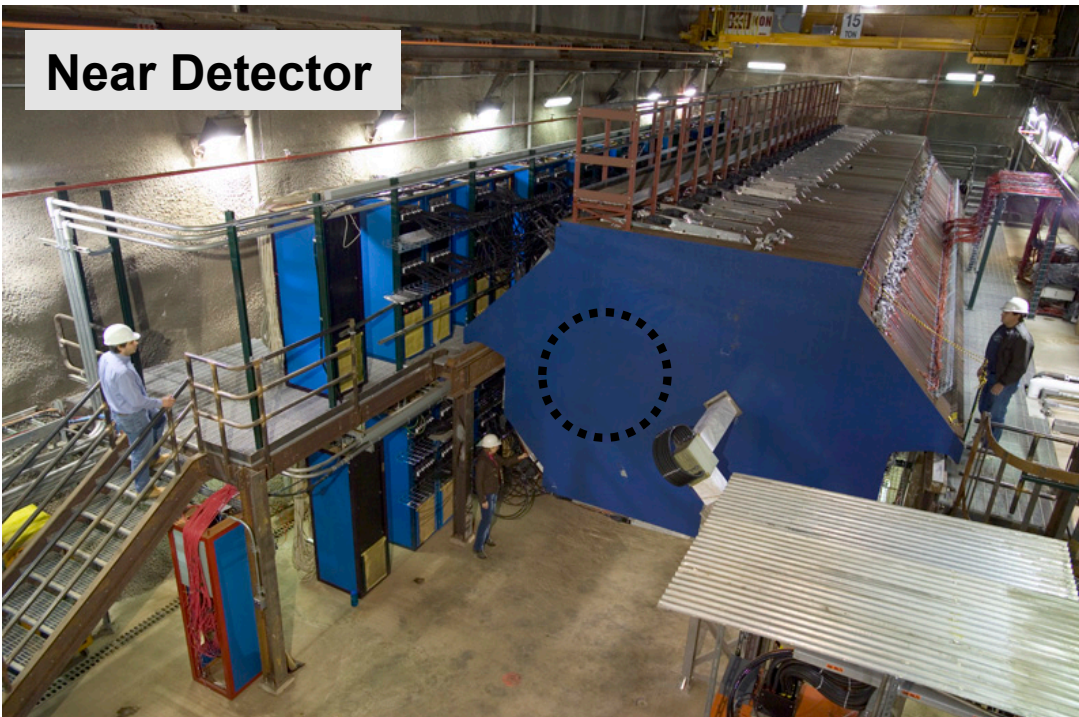


- 120 GeV protons strike target
- 10  $\mu$ s long pulse of  $3 \times 10^{13}$  protons every 2.2 seconds (275 kW)
- Two magnetic horns focus secondary  $\pi/K$ 
  - decay of  $\pi/K$  produce neutrinos
- Variable neutrino beam energy

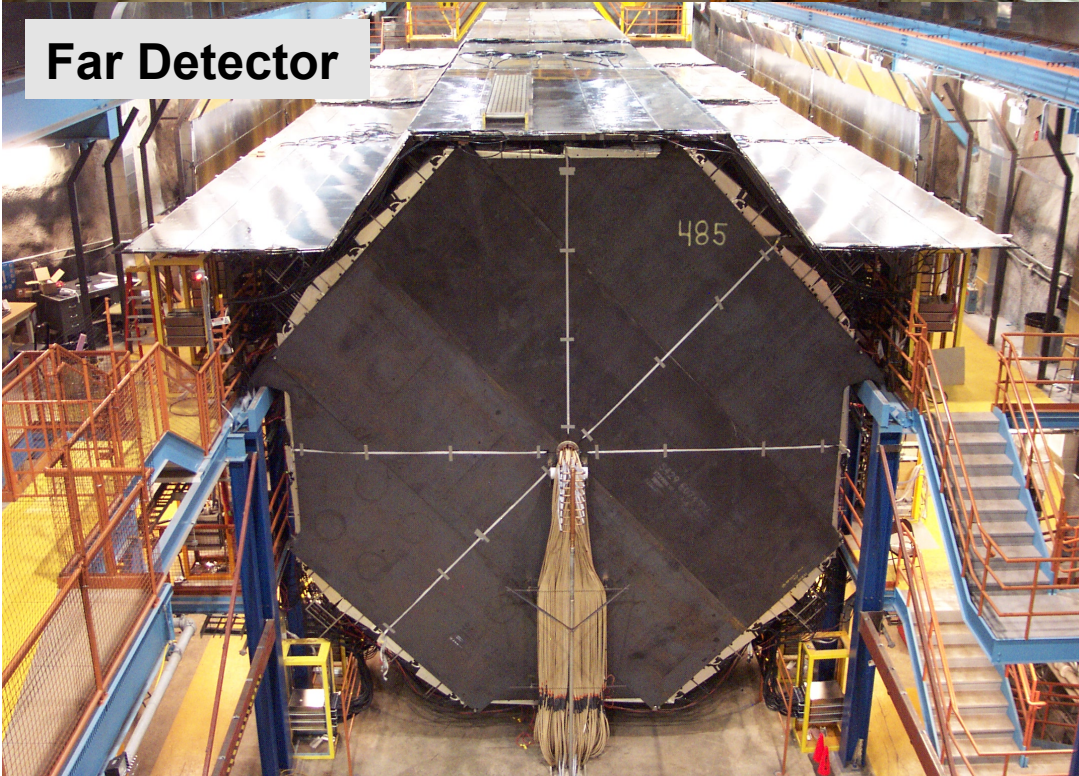




Near Detector



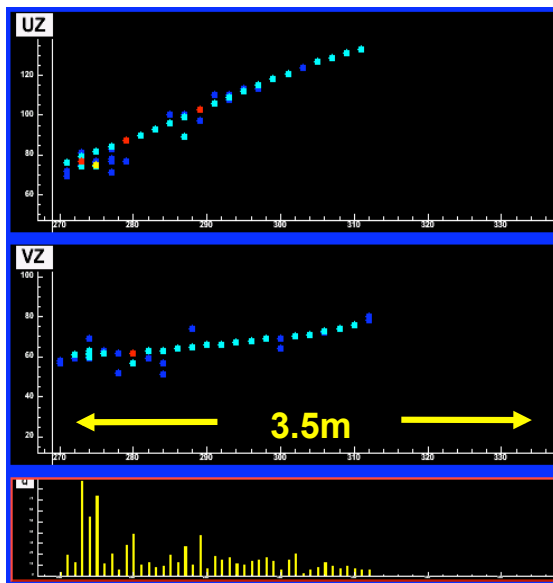
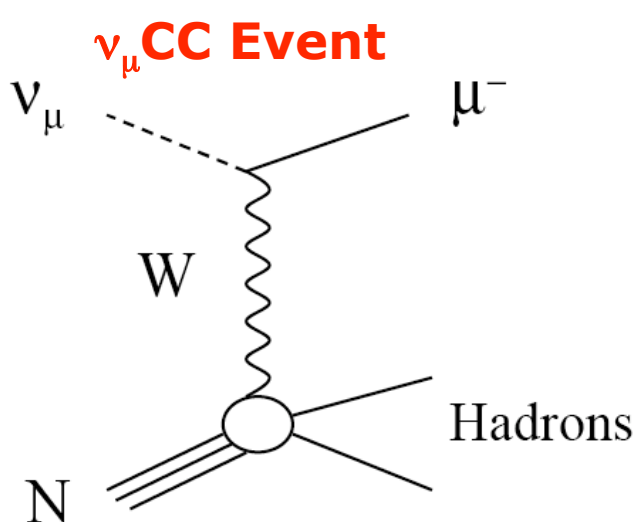
Far Detector



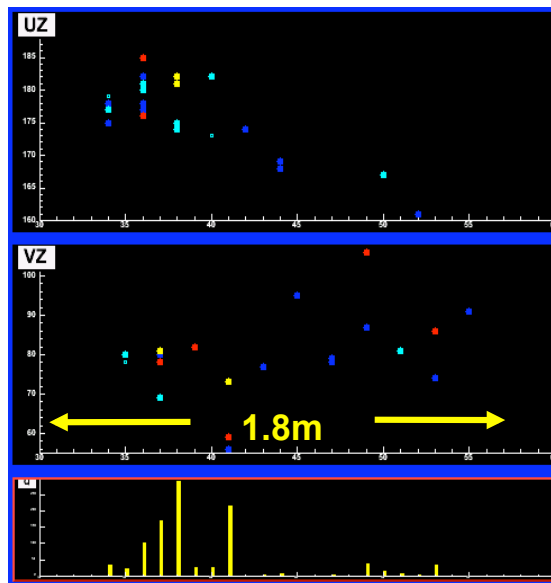
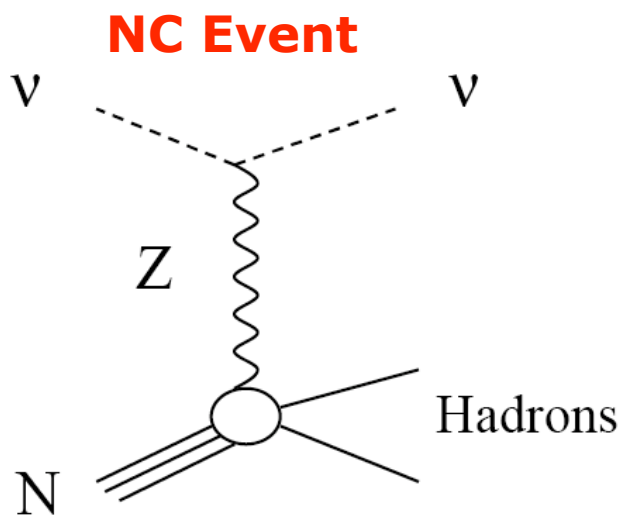
# MINOS Detectors

- Massive
  - 1 kt Near detector
  - 5.4 kt Far detector
- Similar as possible
  - steel planes
    - 2.5 cm thick
  - scintillator strips
    - 1 cm thick
    - 4.1 cm wide
  - Wavelength shifting fibre optic readout
  - Multi-anode PMTs
  - Magnetised ( $\sim 1.3$  T)

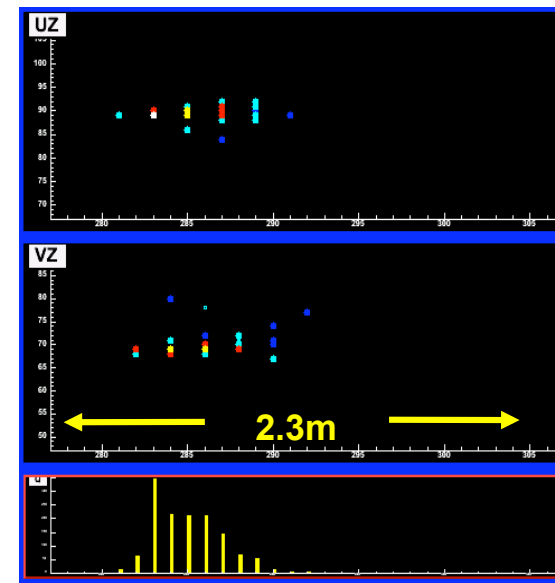
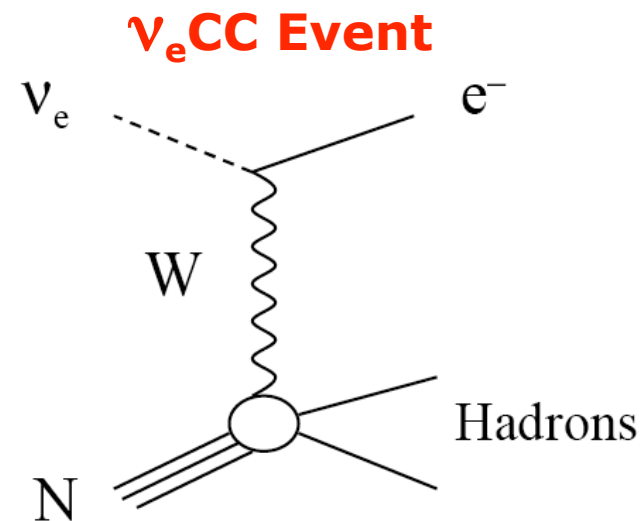
# MINOS Event Topologies (MC)



**long  $\mu$  track+ hadronic activity at vertex**



**short event, often diffuse**



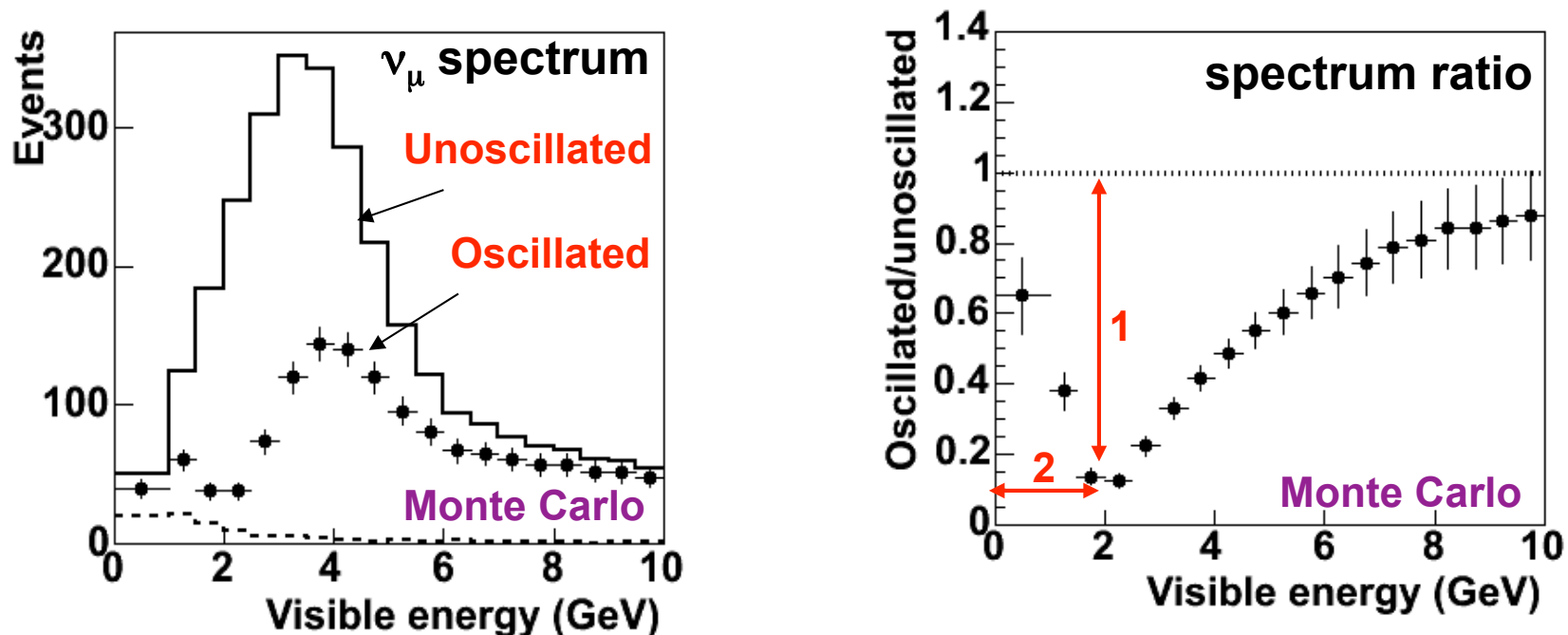
**short, with typical EM shower profile**



# Muon Neutrino Disappearance Analysis

# Experimental Approach

- **Two detector experiment** to reduce systematic errors:
  - Flux, cross-section and detector uncertainties minimised
  - Measure unoscillated  $\nu_\mu$  spectrum at Near detector
    - extrapolate using MC
  - Compare to measured spectrum at Far detector

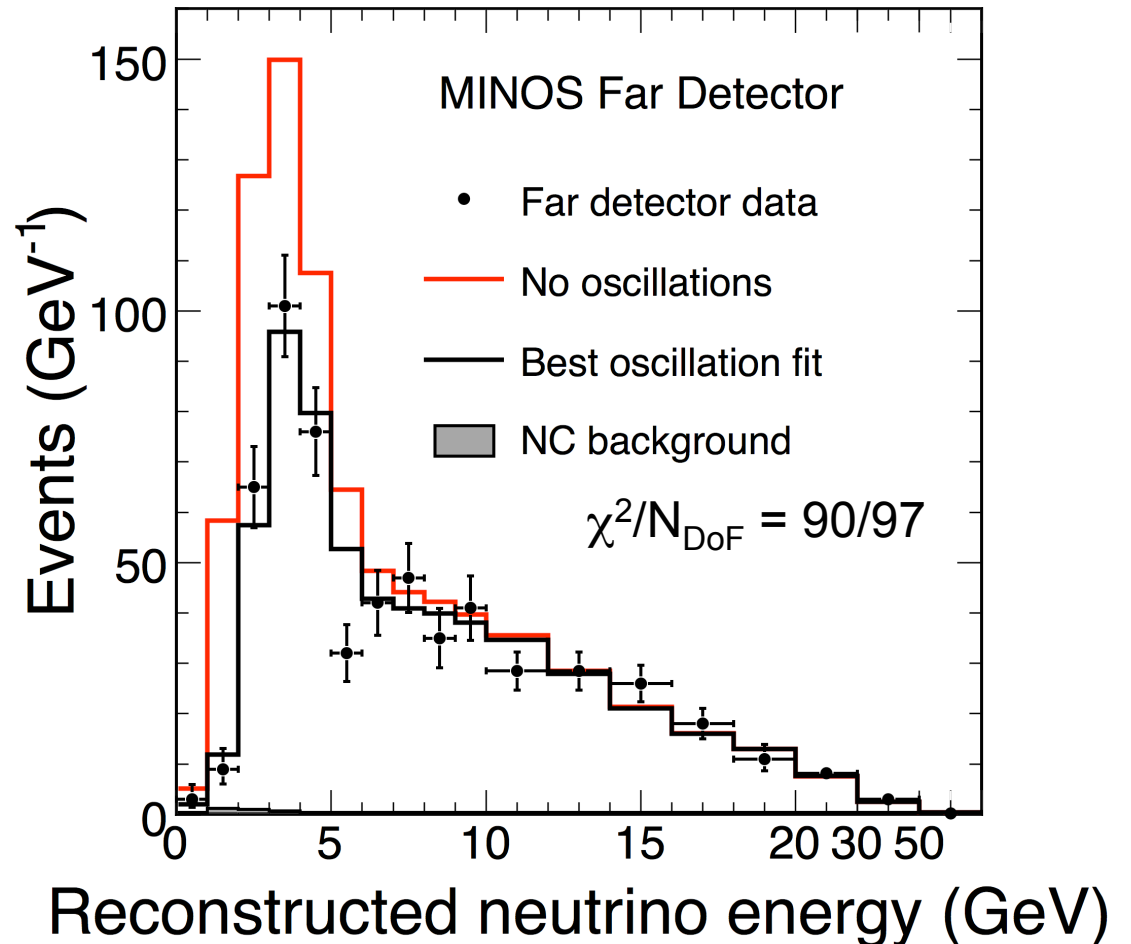


$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \underbrace{\sin^2 2\theta}_1 \sin^2(1.267 \underbrace{\Delta m^2}_2 L / E)$$

# Far Detector $\nu_\mu$ CC Data

- See strong energy dependent distortion of spectrum
- Energy spectrum fit with the oscillation hypothesis:

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2(2\theta) \sin^2\left(\frac{1.27\Delta m^2 L}{E}\right)$$





# Allowed Region

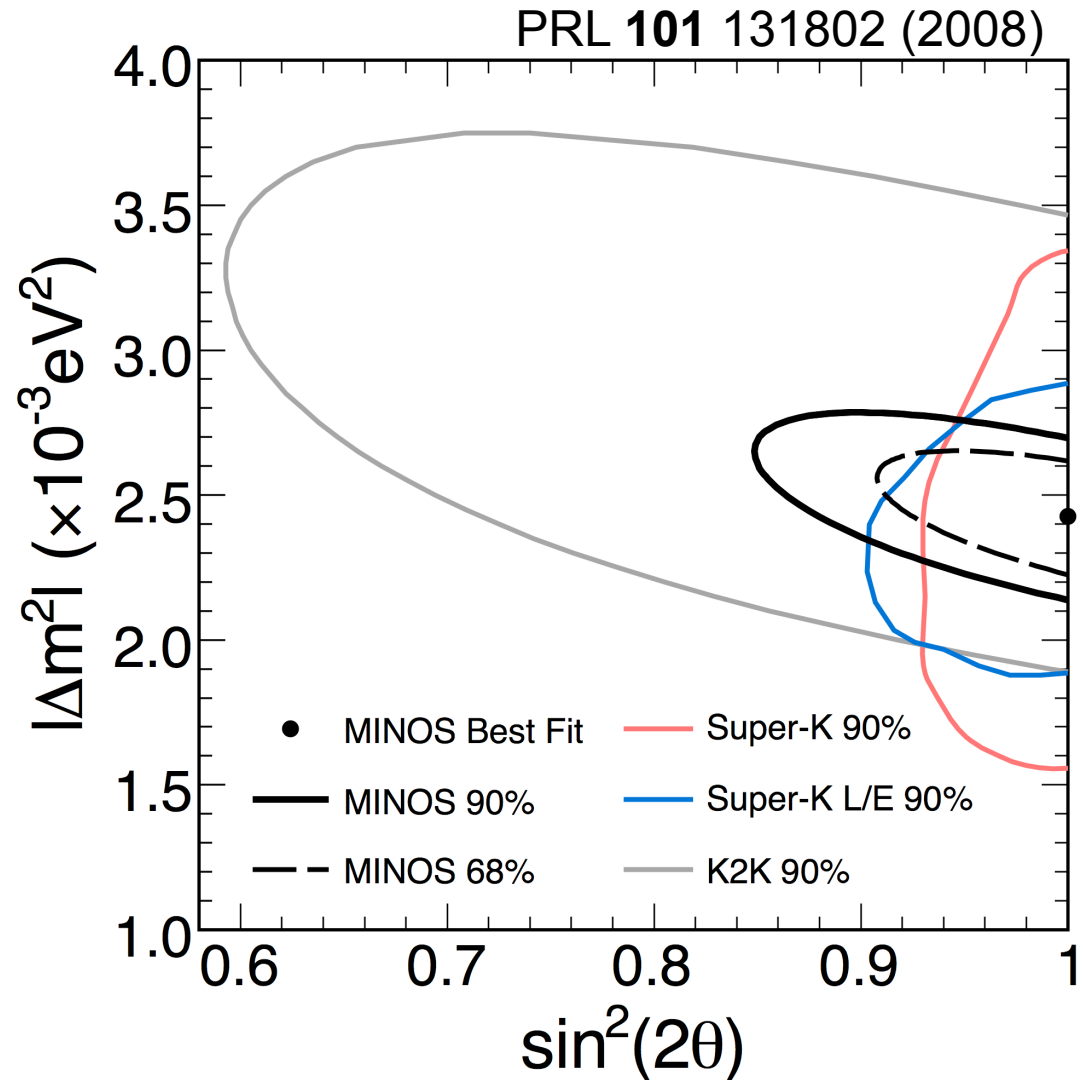
- Fit constrained to physical region and includes 3 largest systematic uncertainties
- Results:

$$|\Delta m^2_{32}| = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$$

at 68% C.L.

$$\sin^2(2\theta_{23}) > 0.90$$

at 90% C.L.



**Most precise measurement of  $|\Delta m^2_{32}|$  performed to date**

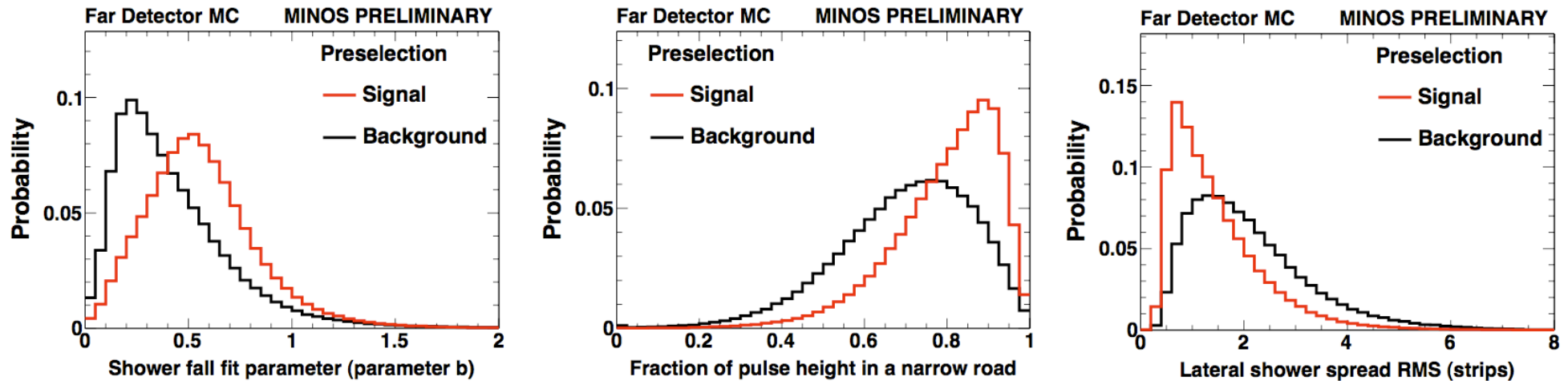
# Electron Neutrino Appearance Analysis (new!)

# $\nu_\mu \rightarrow \nu_e$ Oscillation Search Overview

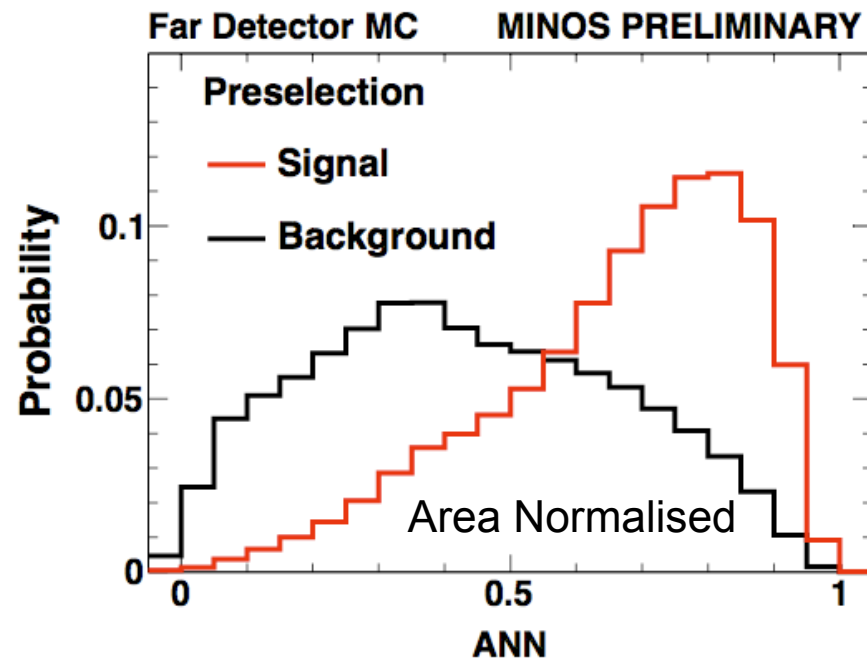
- Sub-dominant neutrino oscillations
  - Look for  $\nu_e$  appearance at Far detector
    - $P(\nu_\mu \rightarrow \nu_e) \approx \sin^2\theta_{23} \sin^2 2\theta_{13} \sin^2(1.27\Delta m^2_{31}L/E)$ 
      - also CPv and matter effects: not shown here but included in fit
  - Electron neutrino events only 2% of total (at Chooz limit)
- Select events w/ compact shower, typical EM profile
  - MINOS optimised for  $\nu_\mu$
  - $\nu_e$  signal selection is harder
    - Steel thickness 2.5 cm = 1.4  $X_0$
    - Strip width 4.1cm  $\sim$  Molière radius (3.7cm)
- Use the Near detector to determine the background



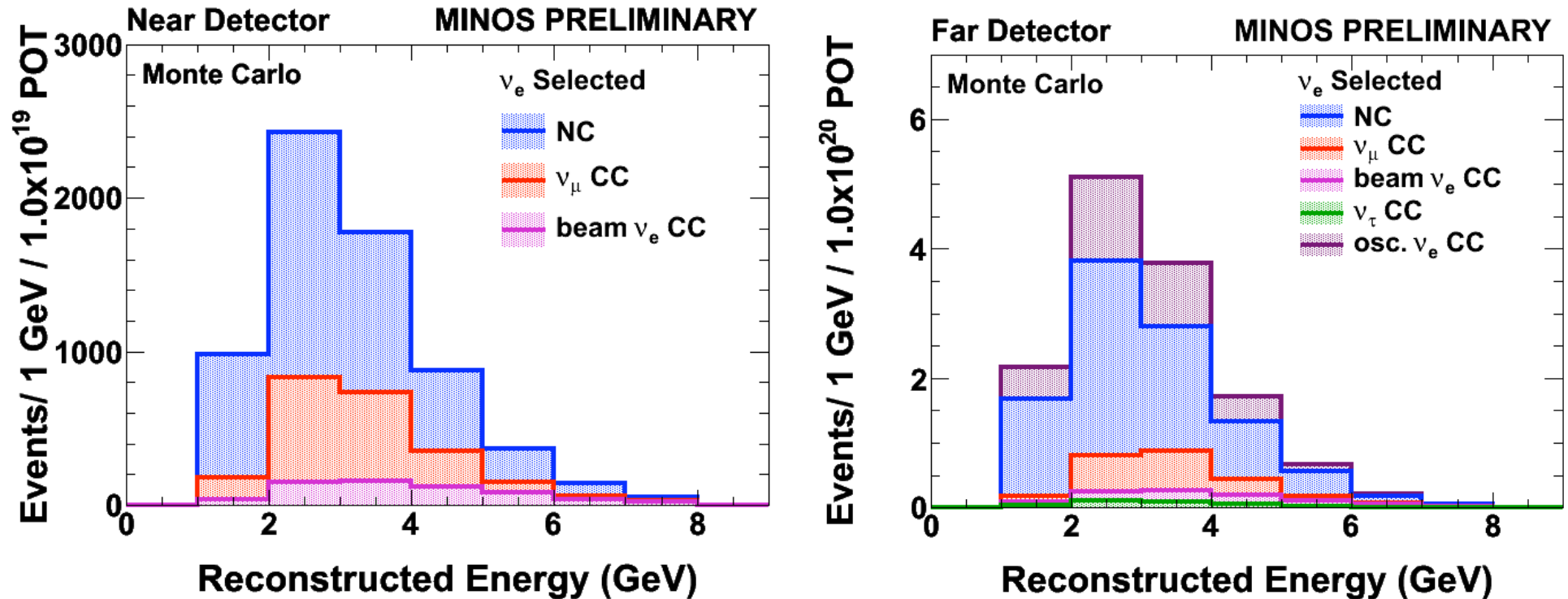
# Selecting $\nu_e$ Events



- 11 variables chosen describing length, width and shower shape
- ANN algorithm achieves:
  - signal efficiency 41%
  - NC rejection >92.3%
  - $\nu_\mu$  CC rejection >99.4%
  - signal/background 1:4 at Chooz limit



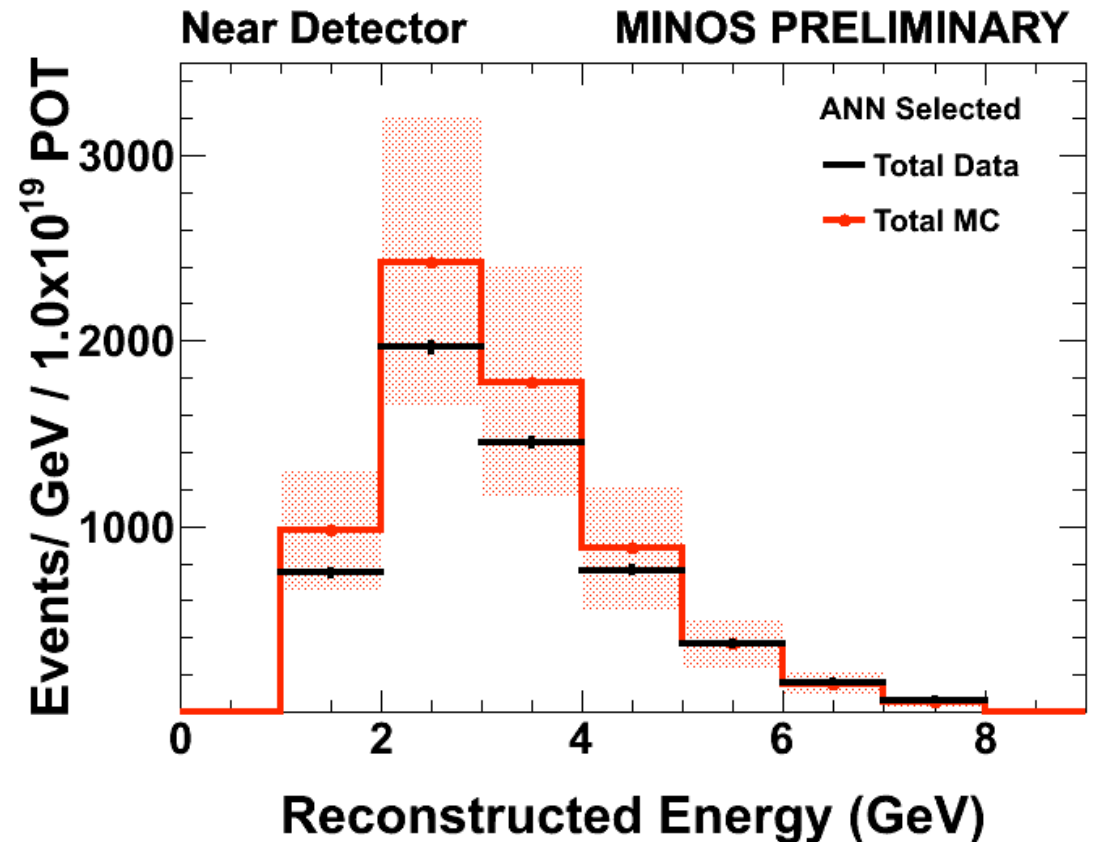
# MC Event Composition in 2 Detectors



- Primary background from NC events, also
  - high- $y$   $\nu_\mu$  CC, beam  $\nu_e$ , oscillated  $\nu_\tau$  at Far detector
- Right plot: purple shows an appearance signal at the Chooz limit ( $\sin^2 2\theta_{13} = 0.15$ )

# Near Detector Data

- $O(20\%)$  data/MC differences in Near det.
  - not surprising, strong background rejection leaves just tails of distributions
  - such differences are expected from the uncertainties in the MC simulation of the hadronisation model
    - tuned to sparse external data
- Use a data driven technique to measure Near detector background
  - compare horn on/off data
    - fit for CC and NC components



**Use measured Near detector background and MC to extrapolate to the Far detector**



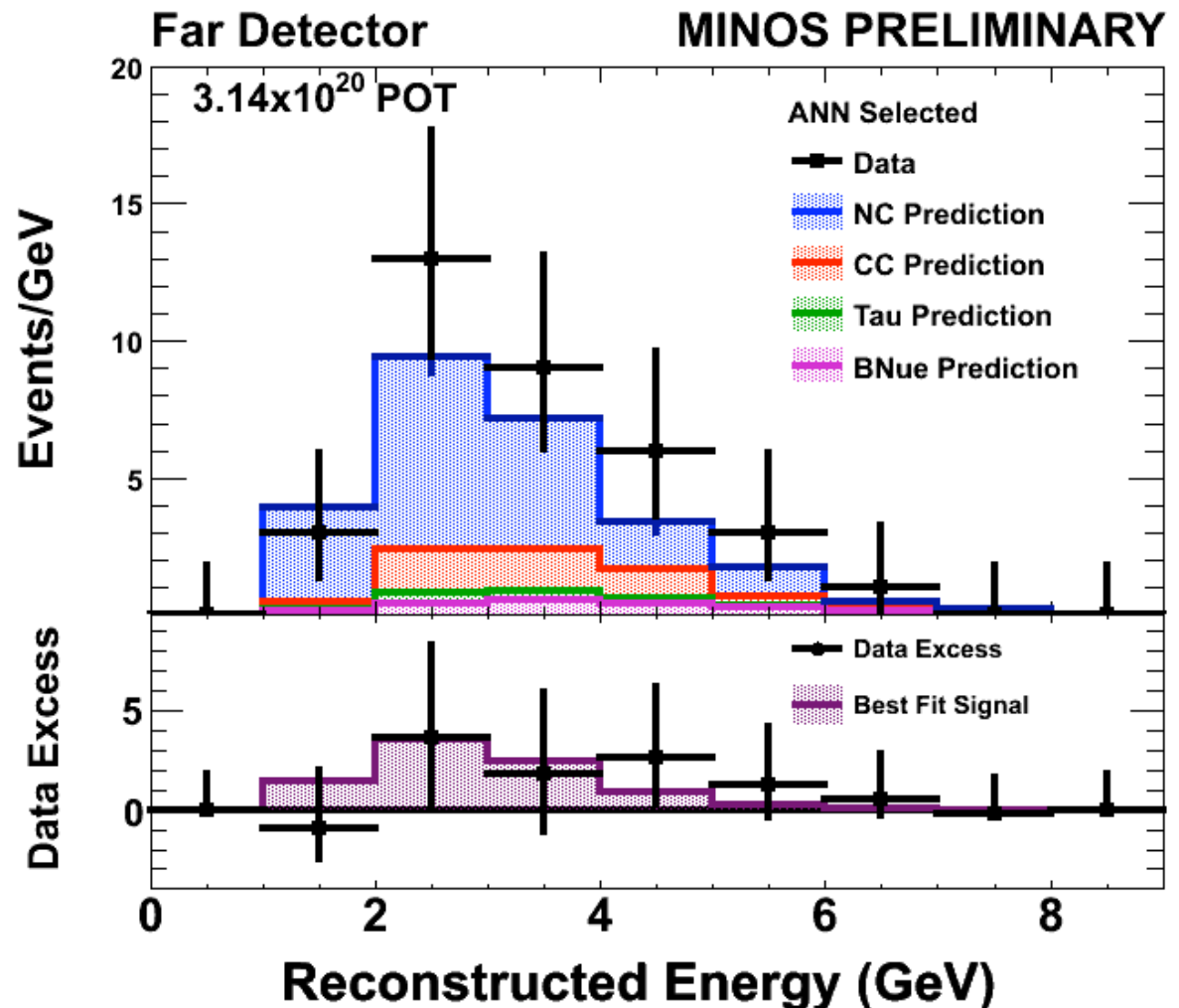
# Preliminary Uncertainties on Background

Preliminary Uncertainties	Size of error
1.) Extrapolation systematic	6.4%
2.) Horn on/off systematic	2.7%
3.) Horn on/off statistical	2.3%
<b>Total (sum in quadrature)</b>	<b>7.3%</b>
<b>Statistical error (data)</b>	<b>19%</b>

- Statistical error dominates
- Systematic error primarily from extrapolating Near detector measurement of background to Far detector

# Far Detector Energy Spectrum

- A blind analysis was performed:
  - all procedures for calculating background and signal were finalised before the Far detector data were looked at
- Expected background:  
 **$27 \pm 5(\text{stat}) \pm 2(\text{sys})$**
- Observed events:  
**35**
- A  $1.5\sigma$  excess over background prediction



Fit the data to the oscillation hypothesis, obtain the signal prediction for the best fit point

# Allowed Region

- A Feldman-Cousins method was used
- Fit simply to the number of events from 1-8 GeV
- Best fit and 90% C.L. limits are shown:
  - for both mass hierarchies
  - at MINOS best fit value for  $\Delta m^2_{32}$  &  $\sin^2(2\theta_{23})$

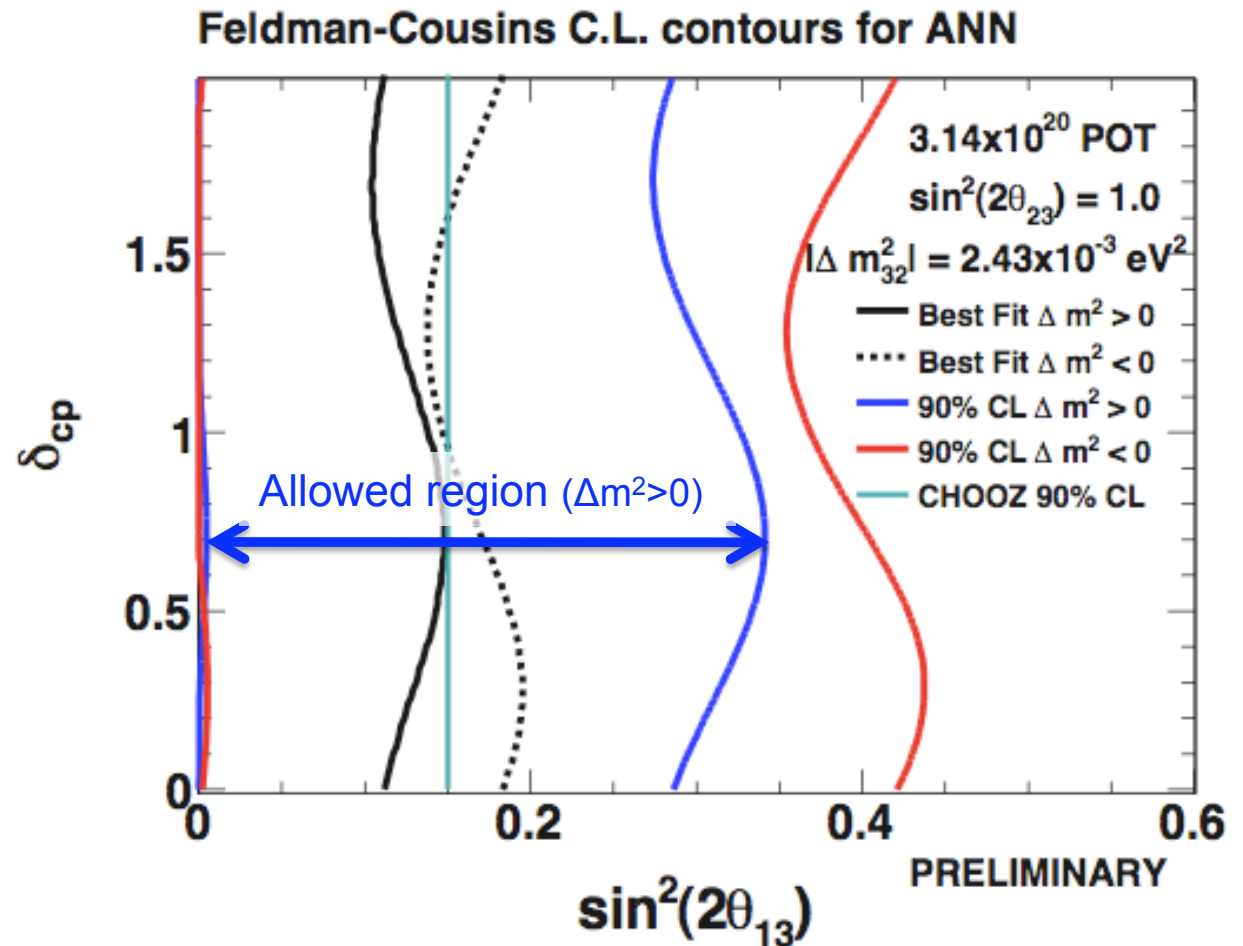
- **Results:**

Normal hierarchy ( $\delta_{CP}=0$ ):

$$\sin^2(2\theta_{13}) < 0.29 \text{ (90\% C.L.)}$$

Inverted hierarchy ( $\delta_{CP}=0$ ):

$$\sin^2(2\theta_{13}) < 0.42 \text{ (90\% C.L.)}$$



# Interpretation

- A 1.5 sigma excess is well within the realms of a statistical fluctuation
- Future  $\nu_e$  appearance analysis:
  - will update MC and reconstruction
  - more than double statistics (already have x2 data!)
  - other analysis improvements too
  - aim to release 2<sup>nd</sup> result a year from now
- Watch this space!



# Neutral Current Analysis

- Looking for sterile neutrino mixing -

# Neutral Current Analysis

- General NC analysis overview:

- All active neutrino flavours participate in NC interaction
- Mixing to a sterile- $\nu$  will cause a deficit of NC events in Far Det.
- Assume one sterile neutrino and that mixing between  $\nu_\mu$ ,  $\nu_s$  and  $\nu_\tau$  occurs at a single  $\Delta m^2$

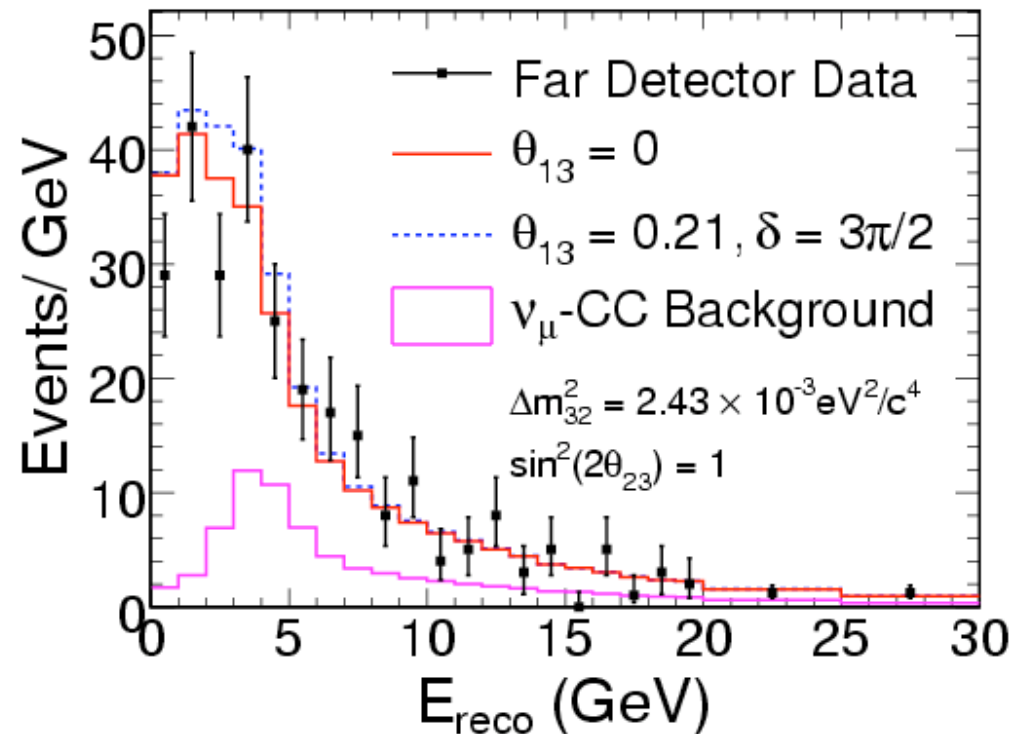
- Survival and sterile oscillation probabilities become:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \alpha_\mu \sin^2(1.27 \Delta m^2 L / E)$$

$$P(\nu_\mu \rightarrow \nu_s) = \alpha_s \sin^2(1.27 \Delta m^2 L / E)$$

( $\alpha_{\mu,s}$  = mixing fractions)

PRL **101** 221804 (2008)



Simultaneous fit to CC and NC energy spectra yields the fraction of  $\nu_\mu$  that oscillate to  $\nu_s$ :

$$f_s = \frac{P(\nu_\mu \rightarrow \nu_s)}{1 - P(\nu_\mu \rightarrow \nu_\mu)} = 0.28^{+0.25}_{-0.28} (\text{stat.} + \text{syst.})$$

$$f_s < 0.68 \quad (90\% \text{ C.L.})$$

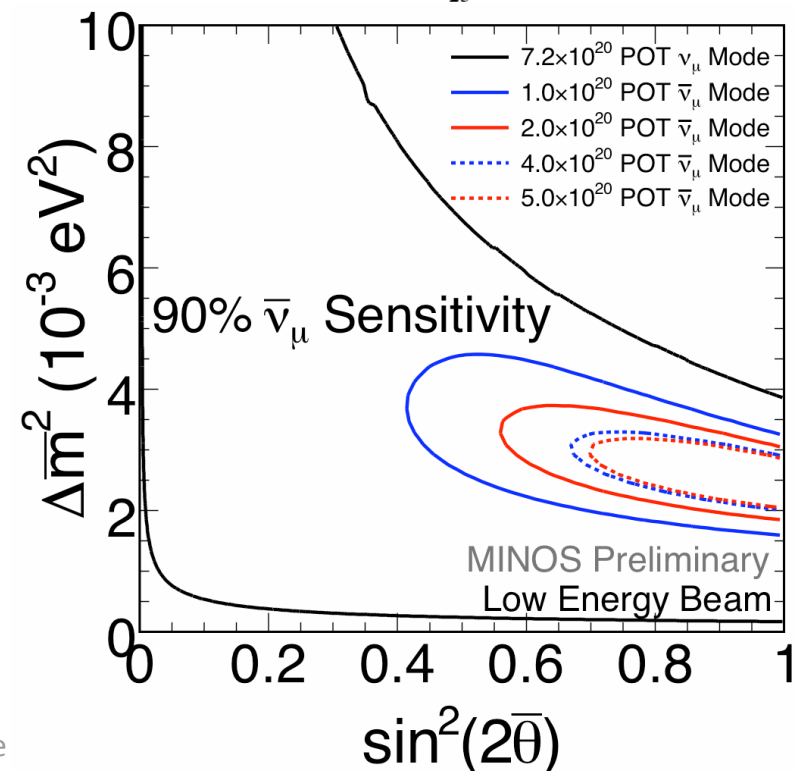
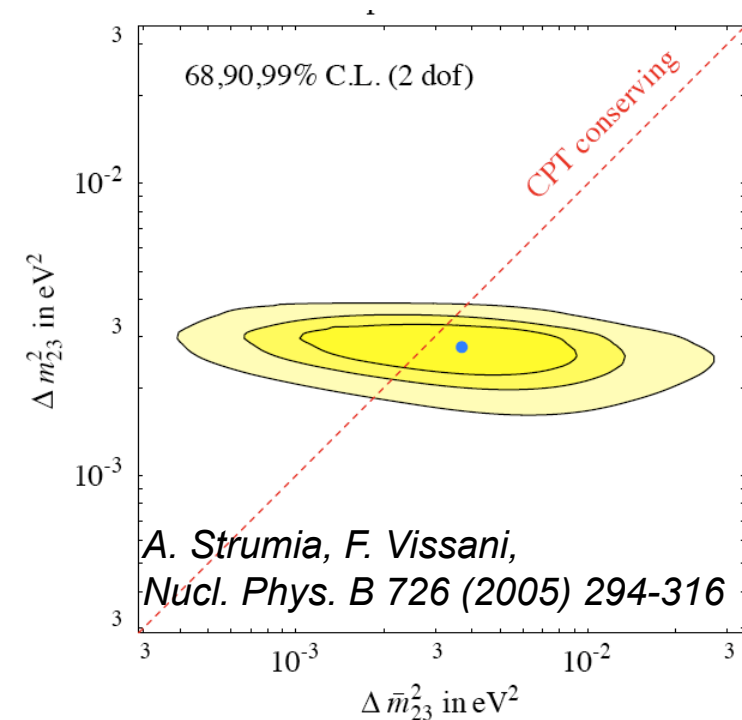
# Future plans

a.) Update all analyses with more than double the data set

b.) Muon antineutrino possibilities

# Muon Anti-neutrino Running

- Neutrino  $\Delta m^2_{32}$  measurement will reach the point of diminishing returns after next result (without accelerator/beam upgrade)
- Possibility to switch beam magnetic horns to focus  $\pi^-$ 
  - create a muon anti-neutrino beam
  - MINOS can make the first direct measurement
    - rapidly reduce the uncertainty on  $\Delta m^2_{32}$  by an order of magnitude



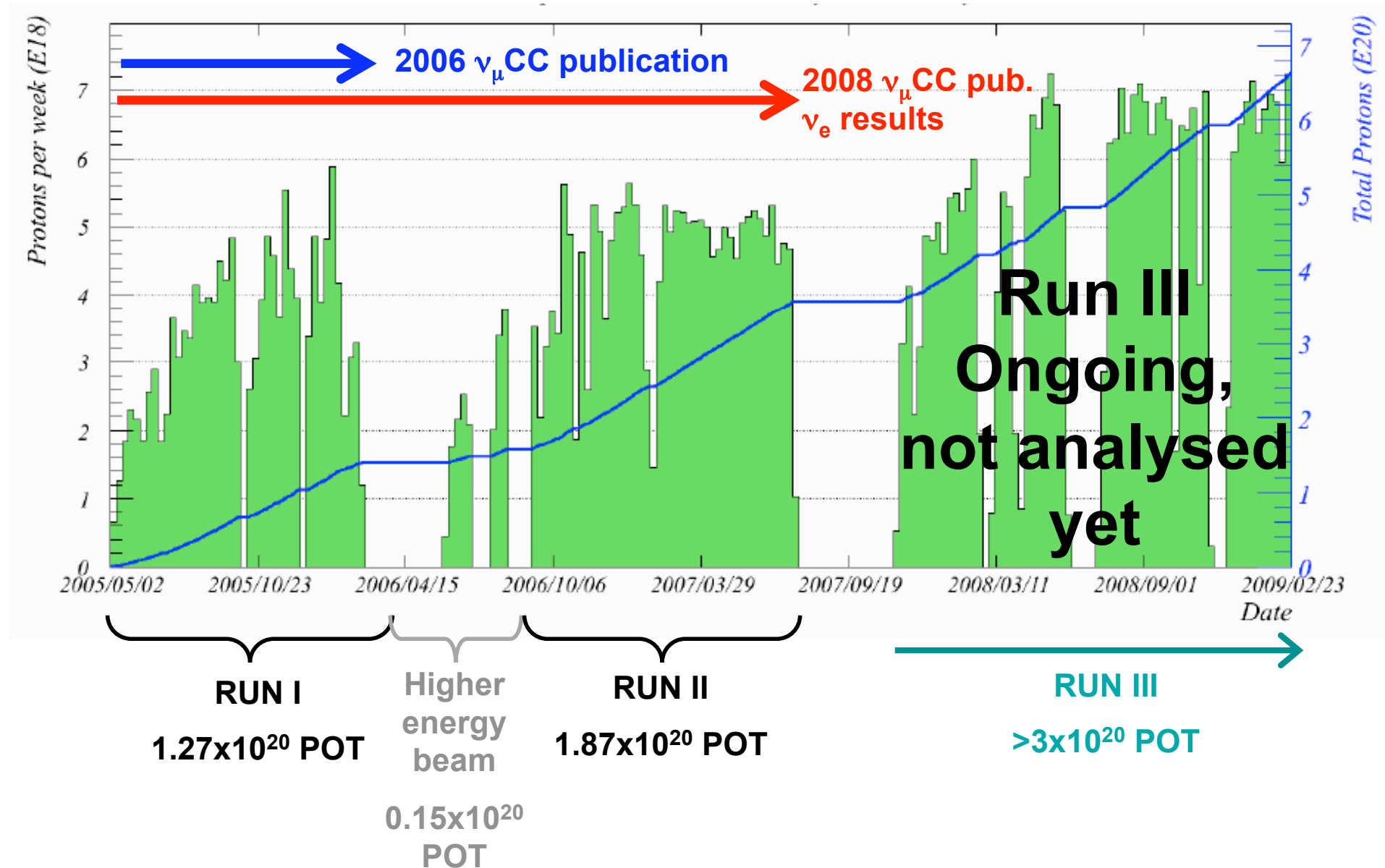
# Conclusions

- MINOS has analysed  $3.2 \times 10^{20}$  POT of beam data ( $>6.6 \times 10^{20}$  POT data now taken)
- Search for electron neutrino appearance
  - $1.5 \sigma$  excess over background prediction
  - $\sin^2(2\theta_{13}) < 0.29$  (90% C.L.) (for normal mass hierarchy,  $\delta_{CP}=0$ )
- Muon neutrino disappearance
  - $|\Delta m^2_{32}| = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$  (68% C.L.)
  - $\sin^2(2\theta_{23}) > 0.90$  (90% C.L.)
- Search for sterile neutrino mixing fraction
  - $f_s < 0.68$  (90% C.L.)
- Muon antineutrino run possibility



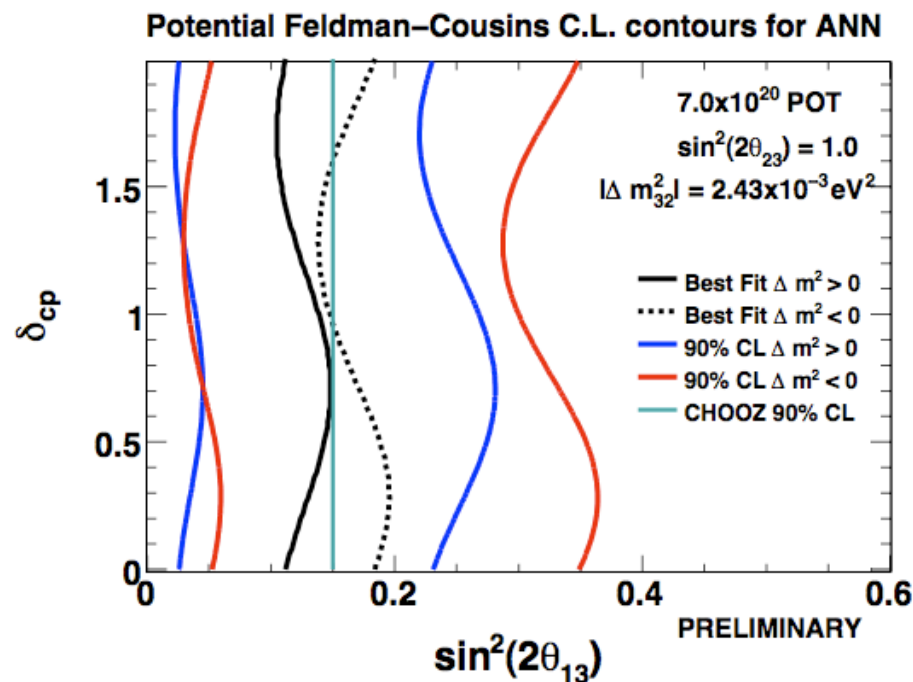
# Backup slides

# Accumulated Beam Data

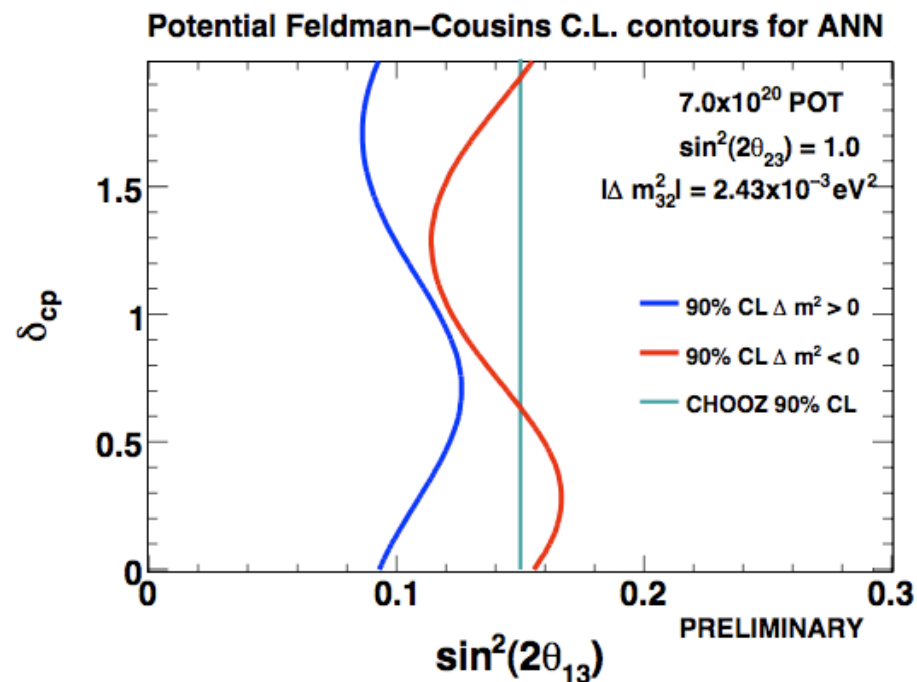


# Future 90% CL contours

**$7.0 \times 10^{20}$  POT**



**Future measurement if data excess persists.**

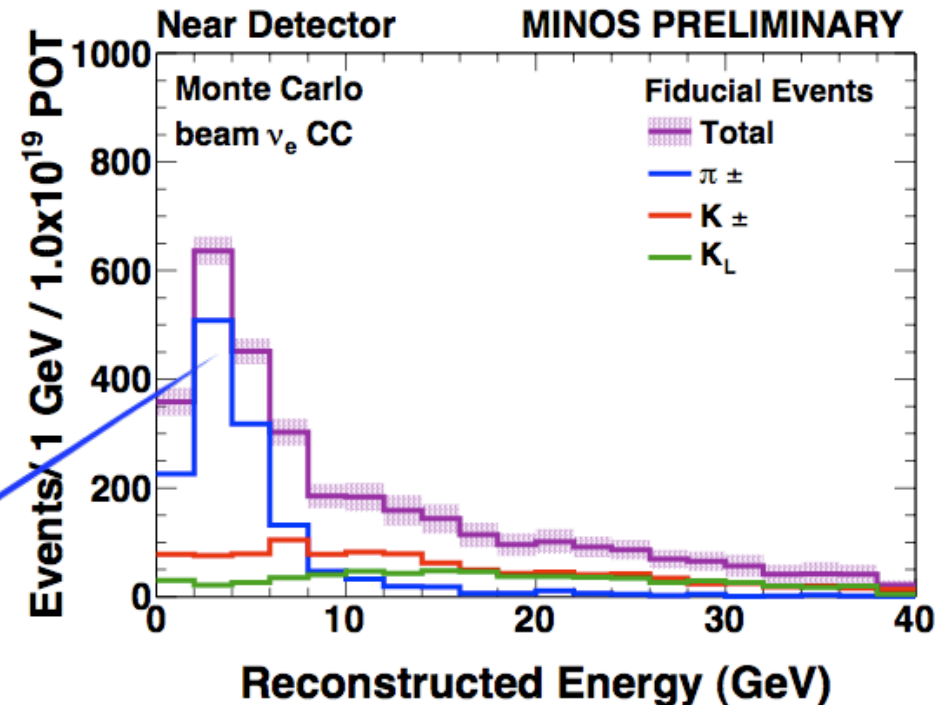
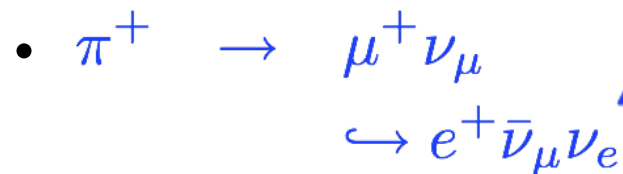


**Future limit if excess cancels with more data.**

**We are close to doubling the data in current running!**

# Beam $\nu_e$ component

- Neutrino beam has 1.3% of  $\nu_e$  contamination from pion and kaon decays.
- Region of interest for the  $\nu_e$  oscillation analysis, 1-8 GeV, dominated by events from secondary muon decays:

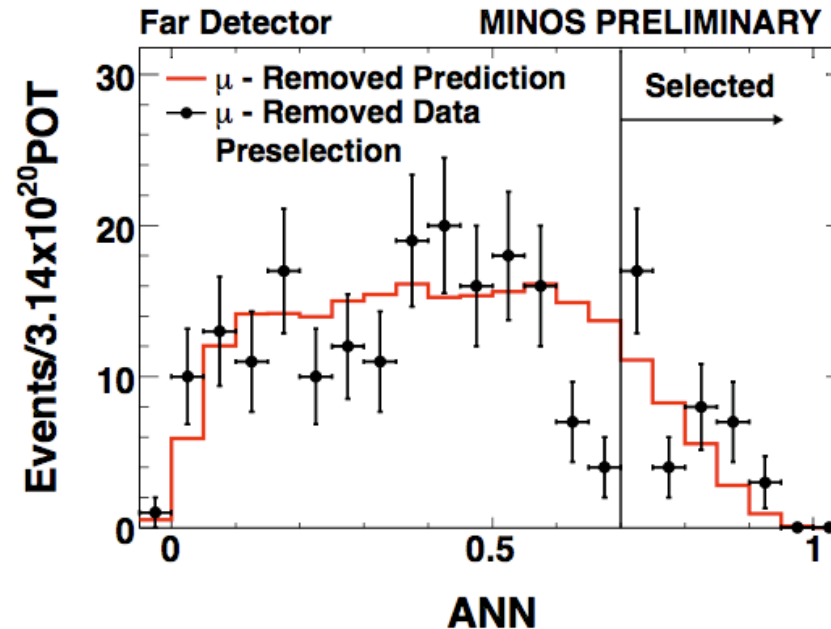


- Near and Far beam  $\nu_e$  spectra are constrained by using  $\nu_\mu$  events from several beam configurations.
- Uncertainties on the flux in the region of interest are  $\sim 10\%$ .

# Sidebands



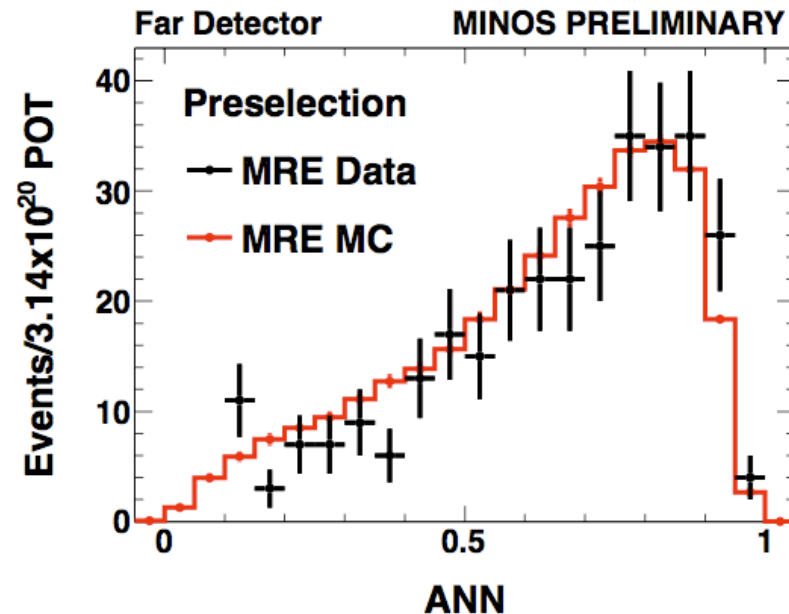
# Muon removed sideband



- We observe a total of 39 events.
- We expect  $29 \pm 5(\text{stat}) \pm 2(\text{sys})$  events.
- Result is within  $2\sigma$
- It is possible this is a statistical fluctuation or it might hint at an unexplained Far/Near difference.

# Muon removed w/electron sideband

- Adding the electron to the muon removed events, present good agreement in PID.

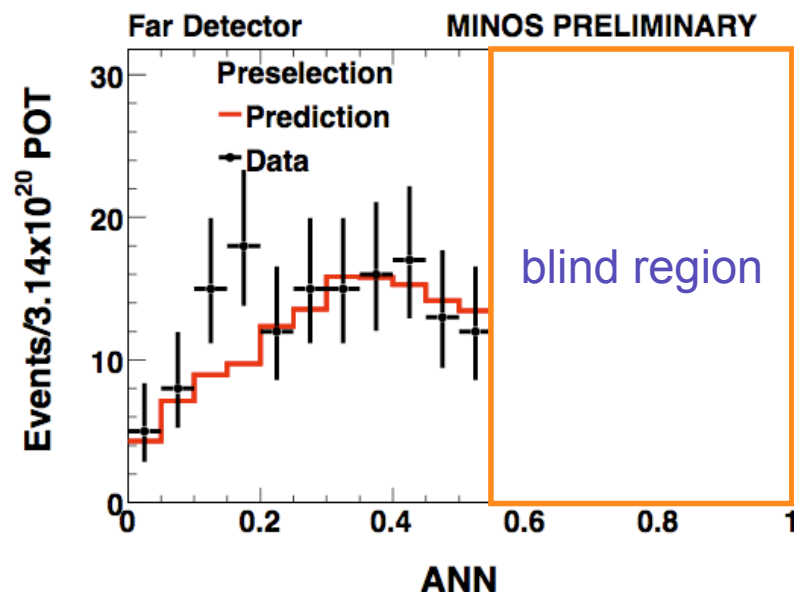


- We observe a total of 159 events.
- We expect  $152 \pm 13(\text{stat}) \pm 12(\text{sys})$  events.

Result is within  $0.5 \sigma$

# Lower PID region

- We also looked in the lower PID region. Finding no obvious disagreement.



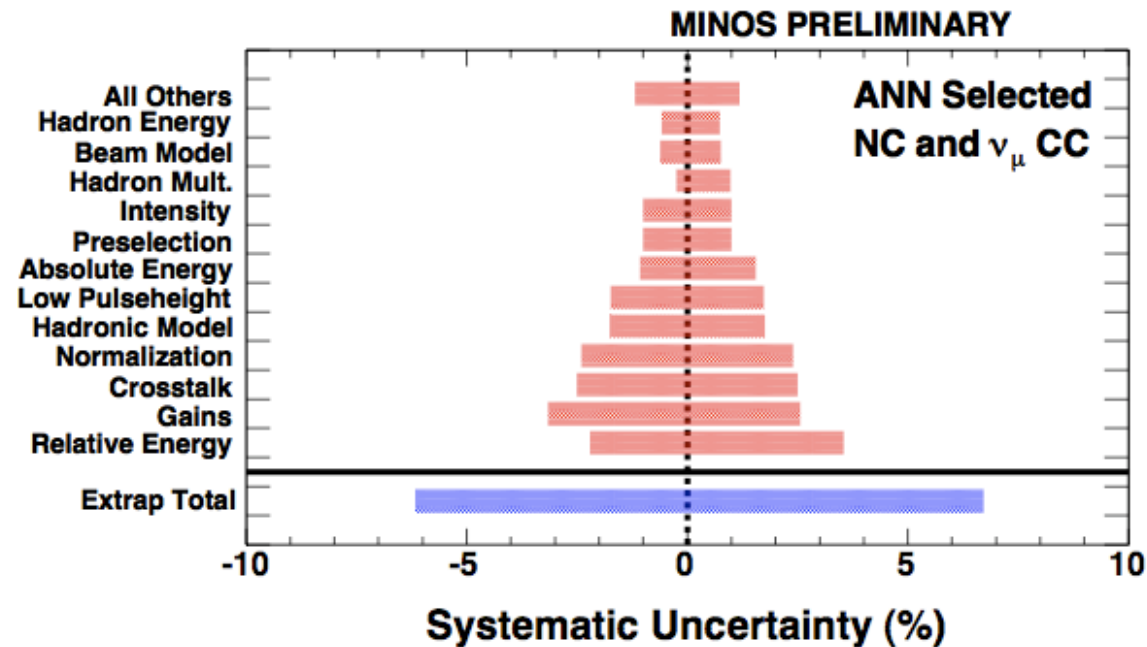
- We observe a total of 146 events.
- We expect  $132 \pm 12(\text{stat}) \pm 8(\text{sys})$  events.

Result is  $\sim 1\sigma$  above expectation

# Systematic Errors

# FD background systematic errors

## Extrapolation errors



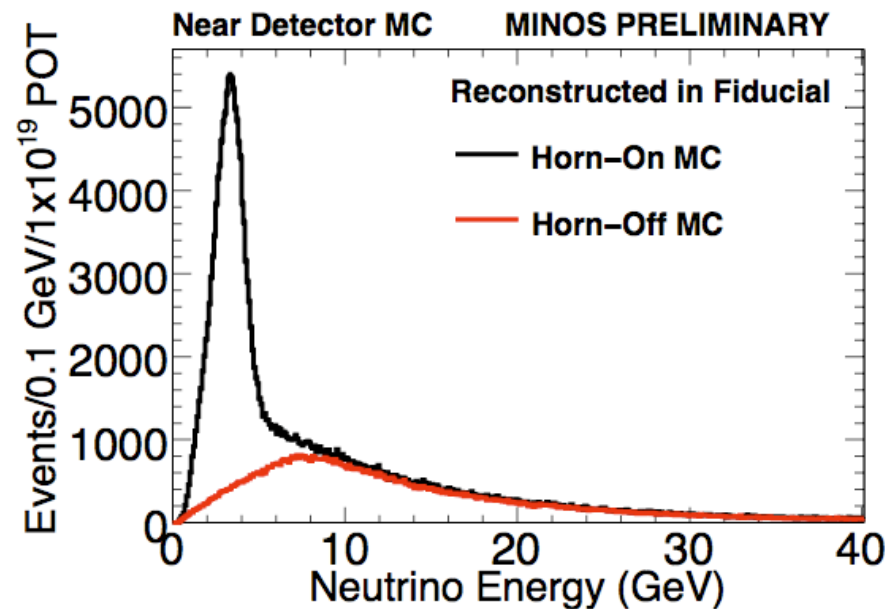
- For most systematic errors, we generated special MC with the modified parameter in Near and Far. Used this modified MC for extrapolation and calculated the difference with the standard results.
- For the main background components the larger systematics are **relative energy, gains, crosstalk and relative normalization**.



# Horn on/off

# Estimating the background using horn on and horn off data

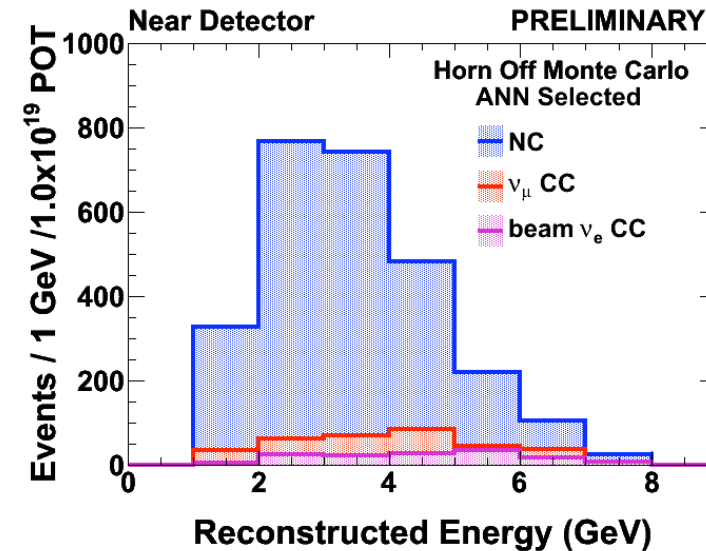
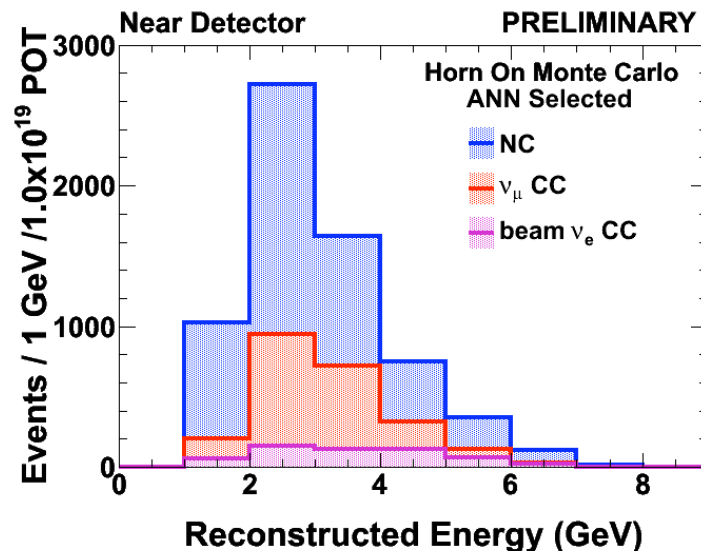
- When beam horns are turned off, the parent pions do not get focused, resulting in the disappearance of the low energy peak in the neutrino energy spectrum.



- The consequence is a spectrum dominated by NC arising from the long tail in true neutrino energy that gets measured in our region of interest in visible energy.

# Estimating the background using horn on and horn off data

- After applying the  $\nu_e$  selection cuts to the ND data, the composition of the selected events is thus very different with the NuMI horns on or off.



- Using the horn off spectrum which is dominated by NC, we can measure that component with better precision than in the horn on beam.

# Estimating the background using horn on and horn off data

- The **beam  $\nu_e$  flux** is obtained from the  $\nu_\mu$  CC flux which is constrained by data in the different beam configurations.
- The two main background components can be estimated using the number of data events in the horn on and horn off configurations:  $N^{\text{on}}$  and  $N^{\text{off}}$ .

$$N^{\text{on}} = N_{\text{NC}} + N_{\text{CC}} + N_e \quad (1)$$

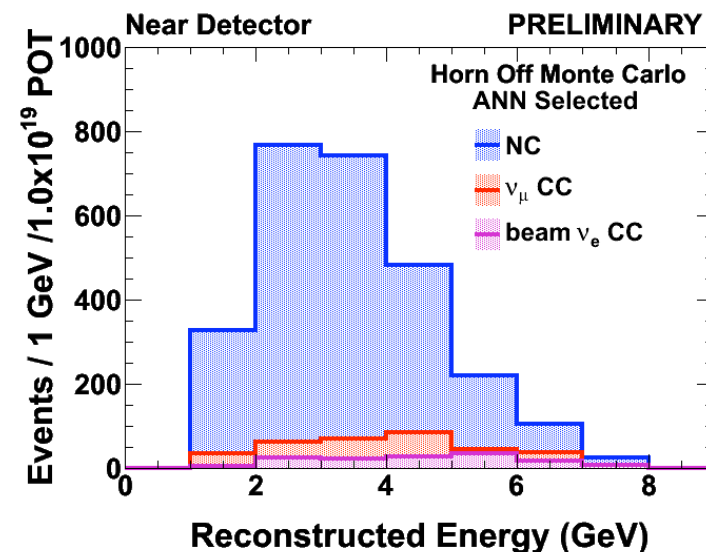
$$N^{\text{off}} = r_{\text{NC}} * N_{\text{NC}} + r_{\text{CC}} * N_{\text{CC}} + r_e * N_e \quad (2)$$

from MC:

$$r_{\text{NC}(\text{CC},e)} = N_{\text{NC}(\text{CC},e)}^{\text{off}} / N_{\text{NC}(\text{CC},e)}$$

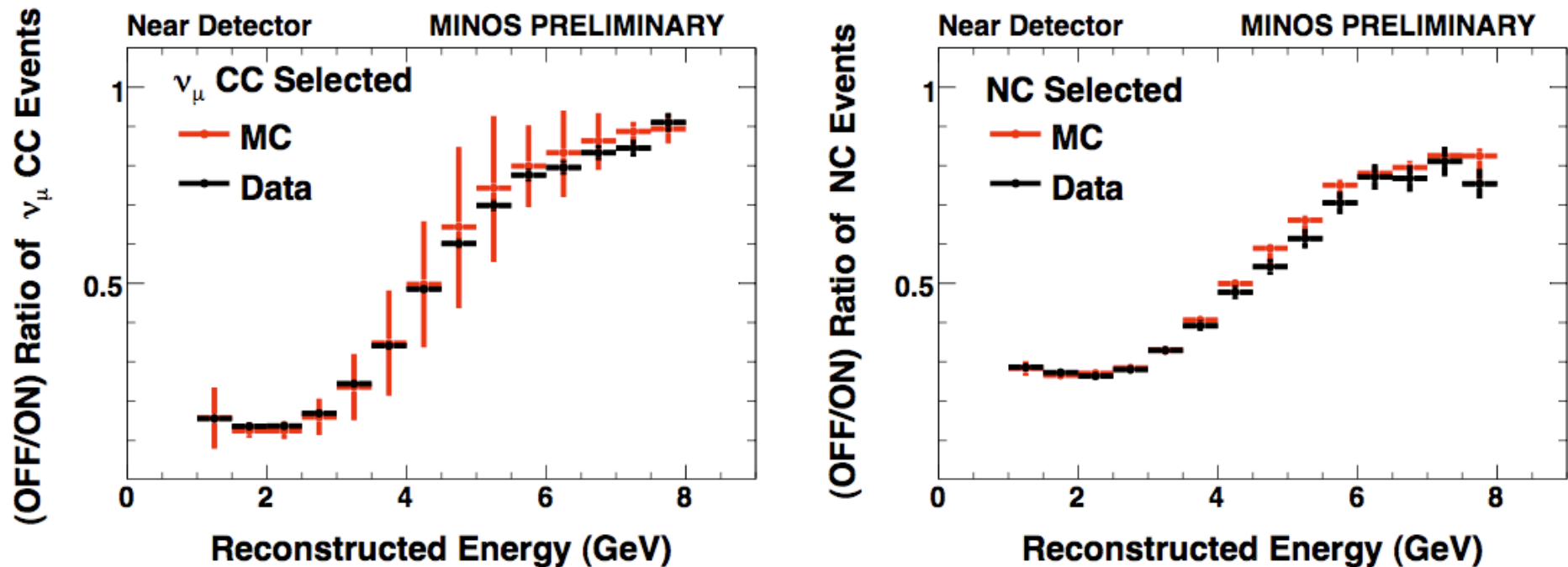
The key is to use the **Horn off/on ratios** for each component to solve:

- Producing **data-driven predictions** for **NC** and  **$\nu_\mu$  CC** background for the horn on configuration.



# Estimating the background using horn on and horn off data

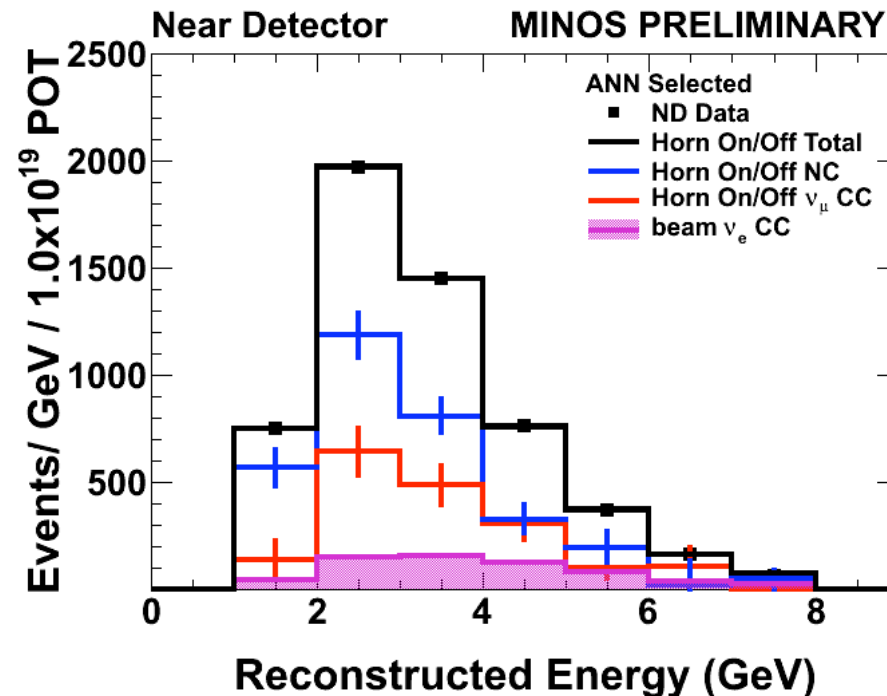
- Horn off/on ratios for  $\nu_\mu$  CC and NC selected events match well between data and MC after fiducial volume cuts.
- Similar ratios are used to solve the horn on/off equations.



MC error statistical plus systematic.

# ND data-driven background

Results from the Horn on/off method

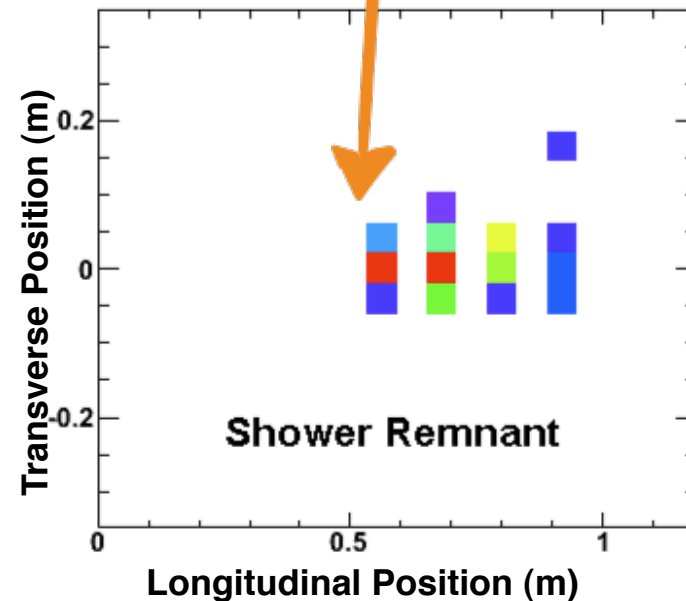
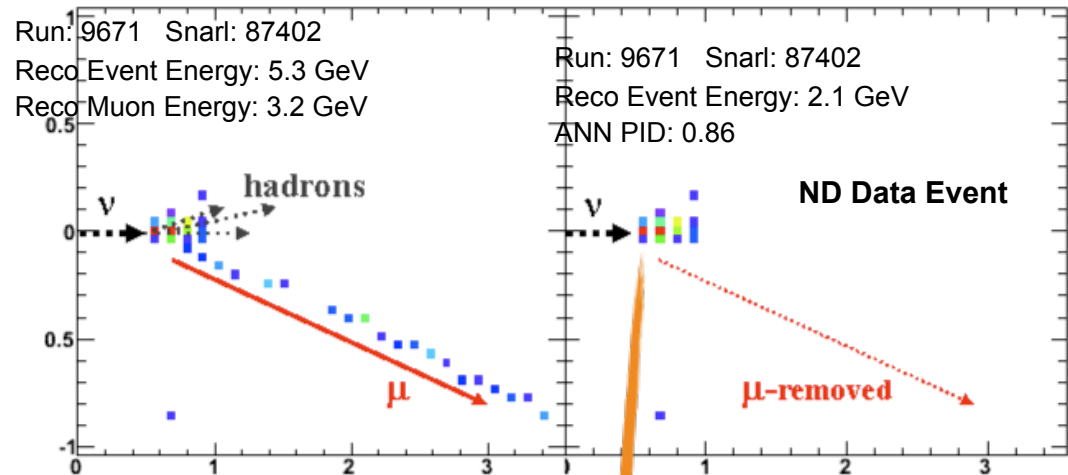


- The NC and  $\nu_\mu$  CC components for the standard beam configuration are simultaneously solved in the horn on/off method and are by definition equal to the data after beam  $\nu_e$  subtraction.

# MRCC

# Studying hadronic showers using muon removal technique

- Remove the muon track in a selected  $\nu_\mu$  CC event and use the rest as a hadronic shower only event.
- We use events that pass our  $\nu_\mu$  Charged Current event selection, i.e. that have a well defined track.
- Well understood  $\nu_\mu$  CC spectra, with well known efficiency and purity from the  $\nu_\mu$  disappearance analysis.

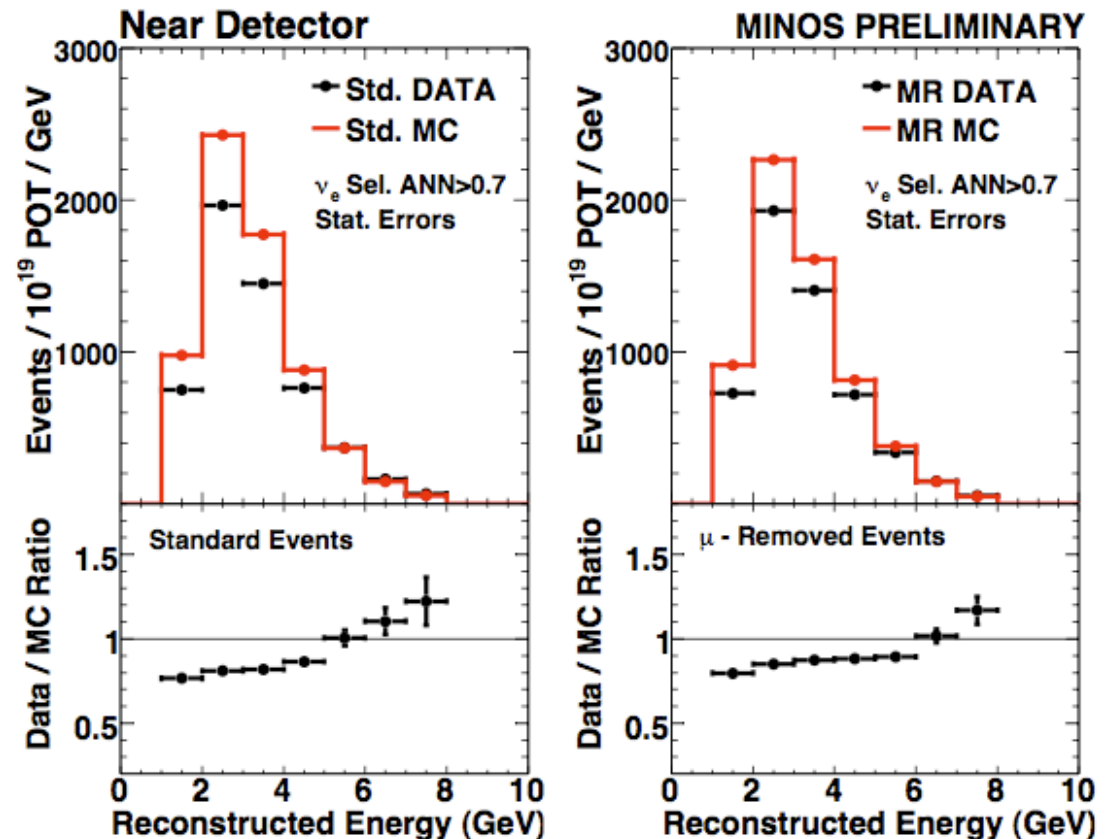


**Muon Removed Charged Current  
events  $\Rightarrow$  MRCC events**



# Hadronic shower modeling in the $\nu_e$ selected data and muon-removed data

- We apply the  $\nu_e$  selection to the standard data and MC as well as to the Muon Removed data and MC.
- Discrepancy with the model shows the same trend not only in energy but in shower topology for both sets.
- Thus modeling of the hadronic shower is a major contribution to the disagreement.
- As the MRCC sample is independent, we can use it to obtain a **data-driven correction** to the model.

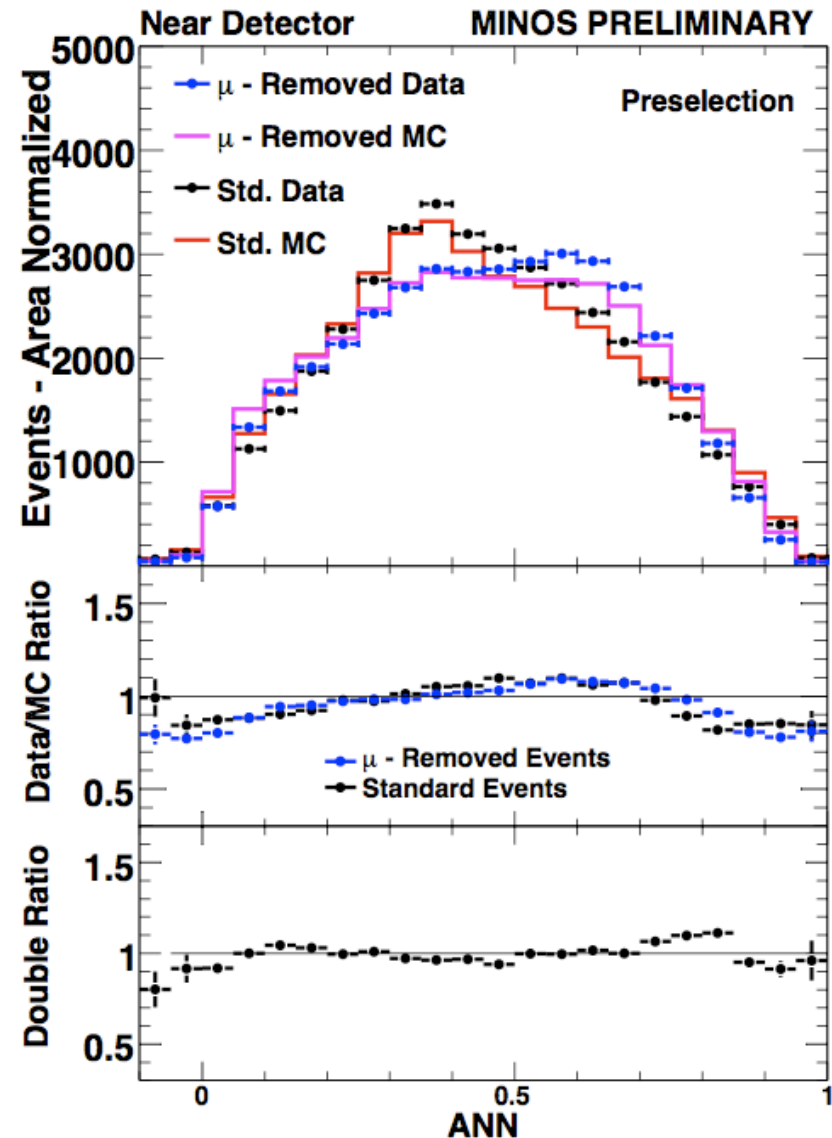


Overall disagreement:

- 16.6% data/MC
- 13.8% MR data/MC

# Hadronic shower modeling in the $\nu_e$ selected data and muon-removed data

- We apply the  $\nu_e$  selection to the standard data and MC as well as to the Muon Removed data and MC.
- Discrepancy with the model shows the same trend not only in energy but in shower topology for both sets.
- Thus modeling of the hadronic shower is a major contribution to the disagreement.
- As the MRCC sample is independent, we can use it to obtain a **data-driven correction** to the model.

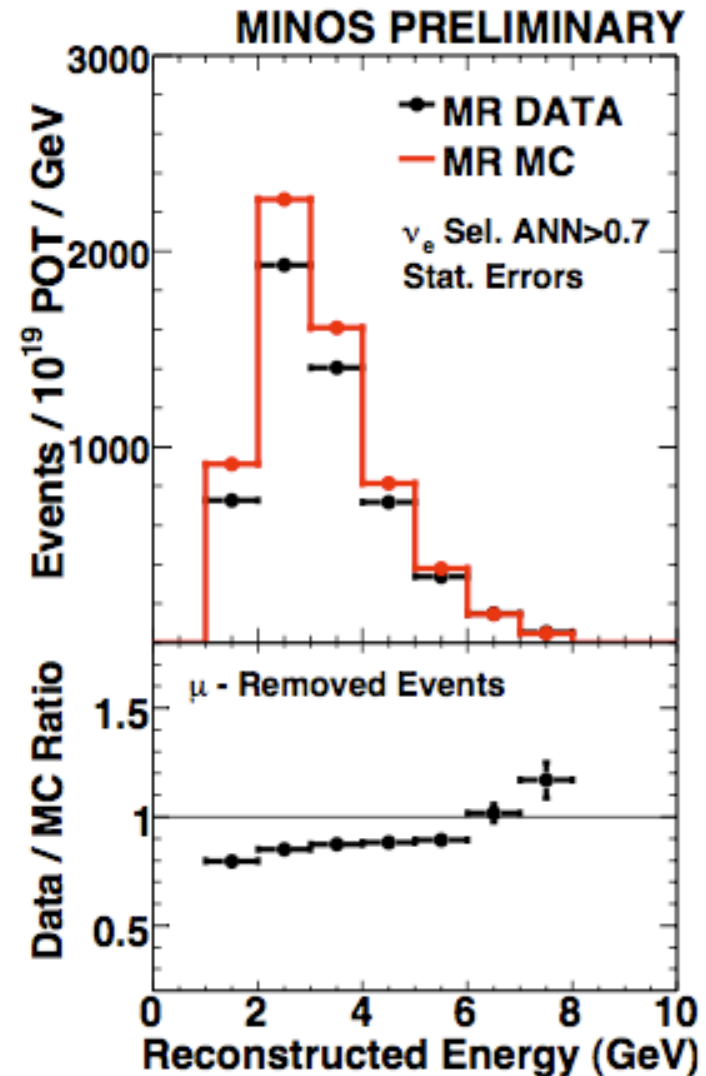


# Using MRCC as a data-driven correction

- We use the data/MC ratio from MRCC to obtain a **data-driven correction** that is applied to the standard NC events as a function of energy.

$$NC_i^{corr} = \frac{MRCC_i^{data}}{MRCC_i^{MC}} \times NC_i^{MC}$$

- The number of  $\nu_\mu$  CC events is taken from the number of events in the data minus the corrected NC and beam  $\nu_e$  events.
- Differences between NC and MRCC showers introduces a systematic error that is difficult to quantify.

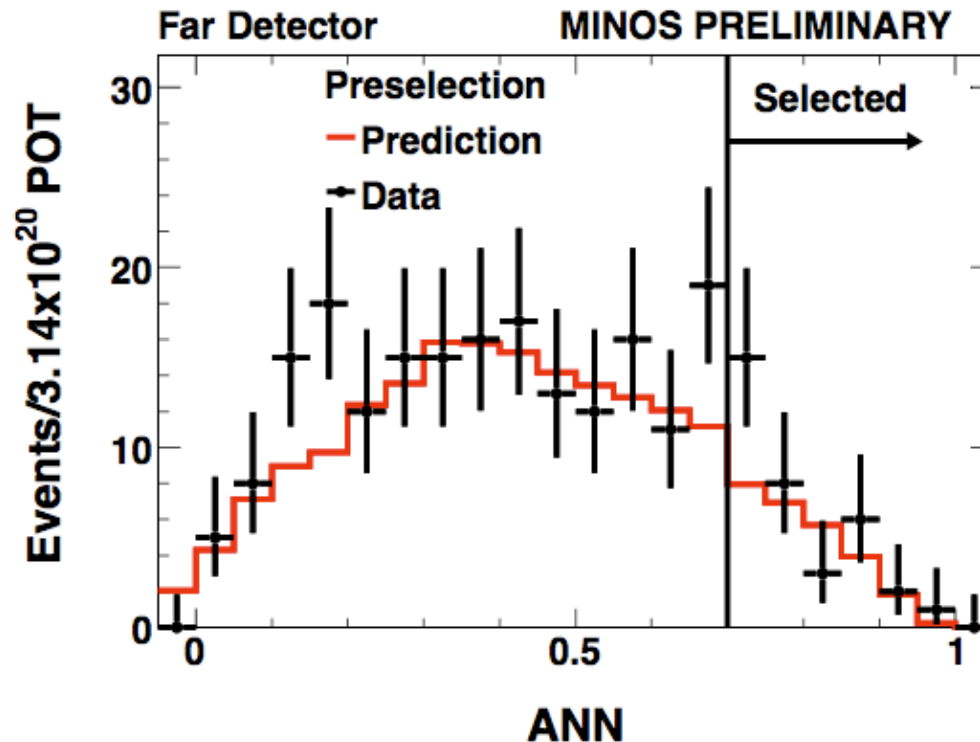


## Secondary separation method

# Far detector selected events

# $\nu_e$ Selected Far Detector Data

- Preselected data in the FD as a function of PID compared to the corrected MC.

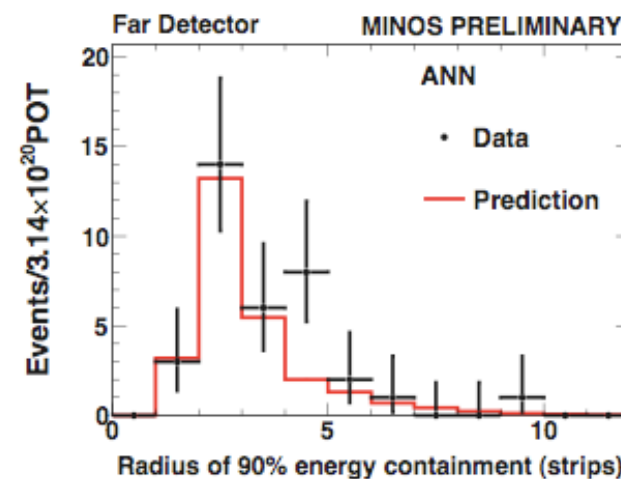
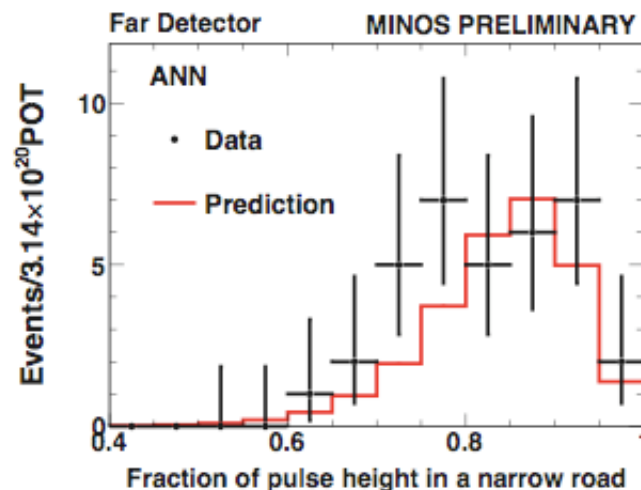
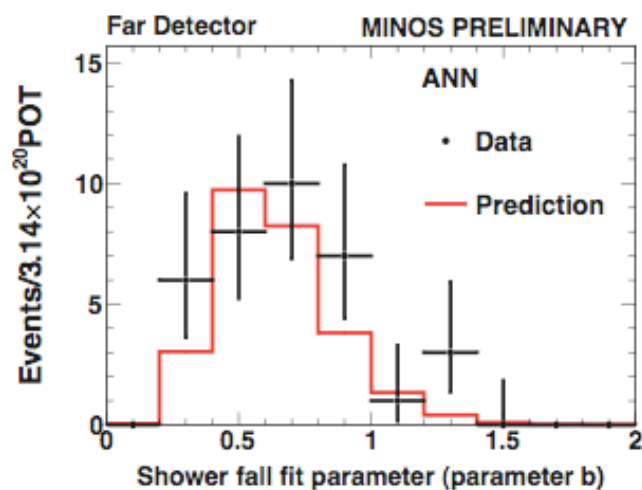
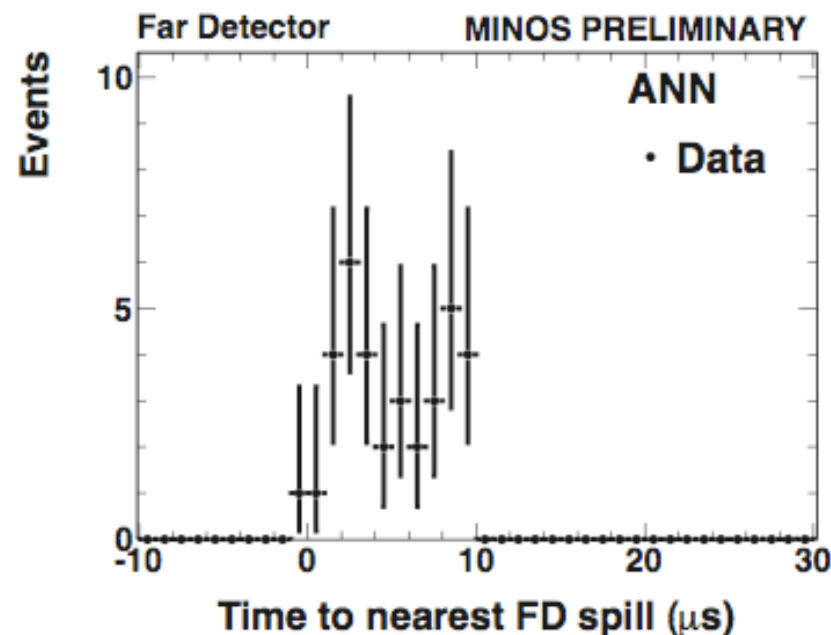
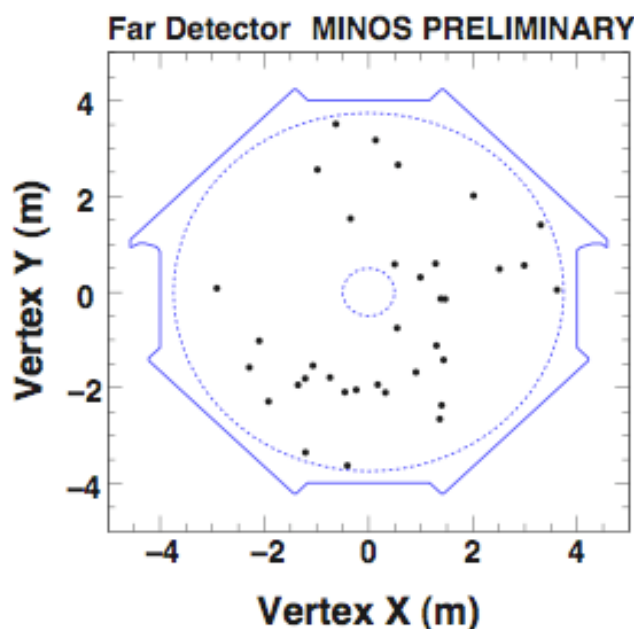


- We observe a total of 35 events.
- We expect  $27 \pm 5(\text{stat}) \pm 2(\text{sys})$  background events.

Results are  $1.5 \sigma$  above expected background.

# Far Data $\nu_e$ Selected Distributions

35 events seen for  $3.14 \times 10^{20}$  POT

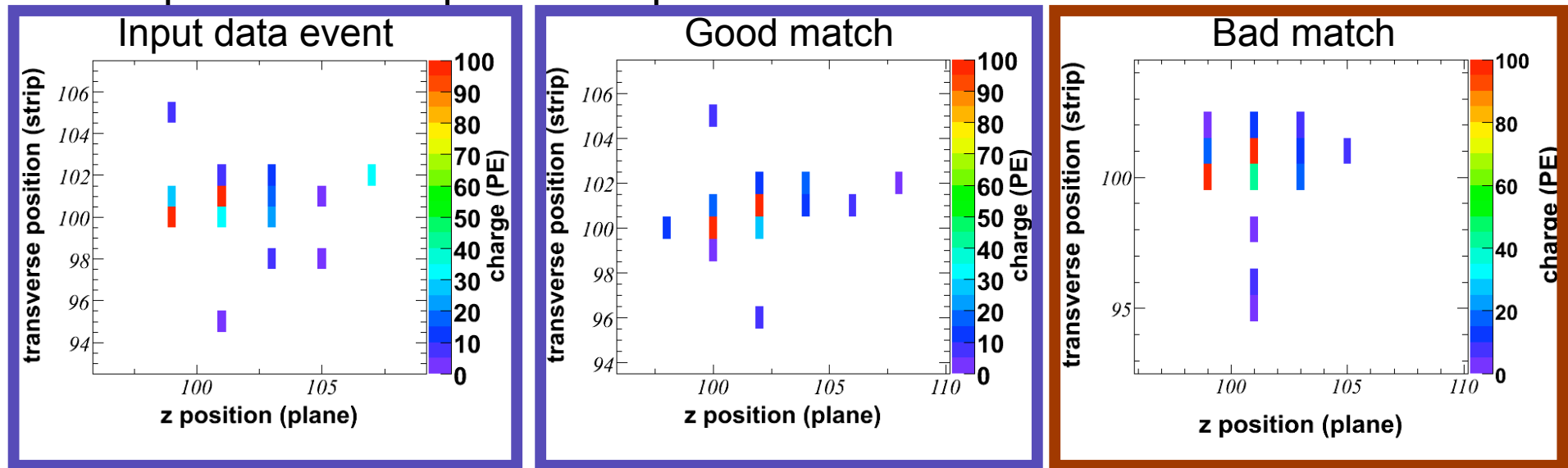


# Library Event Matching

# Alternative selection algorithm

## Library Event Matching (LEM)

What is the likelihood that two events come from the same hit pattern at the photomultiplier level?

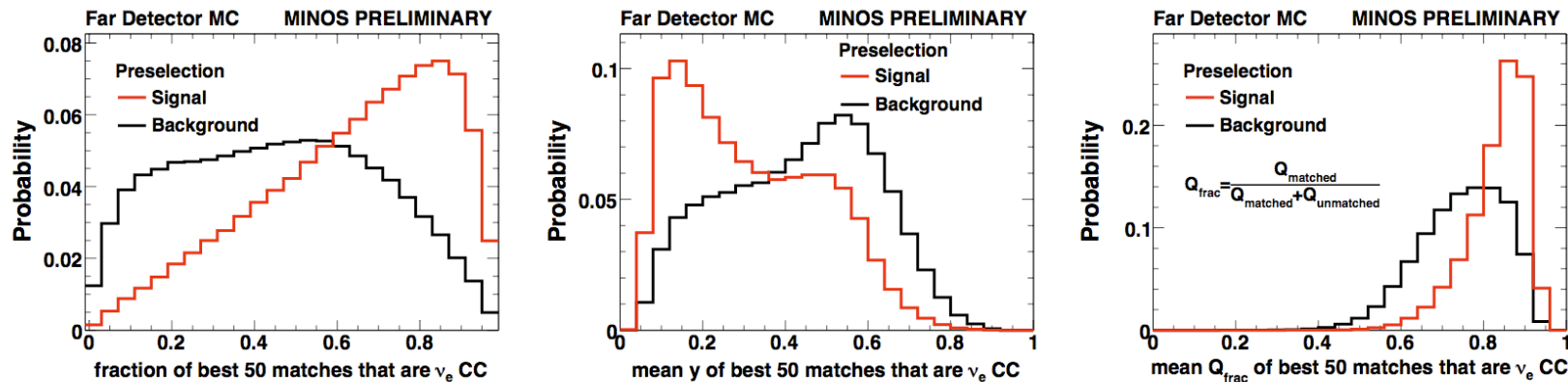


- Compare each input event to large library of MC  $\nu_e$  CC and NC events.
- Select 50 best matches according to the likelihood that two events have the same hit pattern in position and energy deposition.
- Construct discriminant variables from the properties of the 50 best matches, eg. fraction of the 50 best matches that are  $\nu_e$  CC.

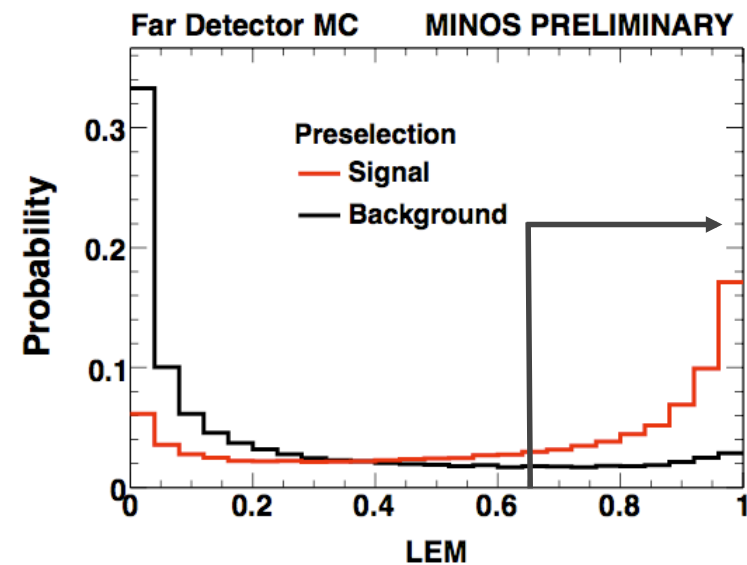


# Selecting $\nu_e$ events with LEM

fraction of electron neutrino events in 50 best matches



- 3 variables combined in a likelihood as a function of energy.
- LEM algorithm has better signal efficiency and background rejection.
- Sidebands may indicate an unexplained Far/Near difference to which this method would be more sensitive.



**Secondary selection method**

Area Normalized

$$\Delta m_{32}^2 = 0.0024 \text{ eV}^2, \sin^2 \theta_{23} = 1.0$$

# CC Analysis

# Alternative Models

Two alternative disappearance models are disfavoured

**Decay:**

$$P_{\mu\mu} = \left( \sin^2(\theta) + \cos^2(\theta) \exp(-\alpha L/2E) \right)^2$$

V. Barger *et al.*, PRL82:2640(1999)

$$\chi^2/\text{ndof} = 104/97$$

$$\Delta\chi^2 = 14$$

**disfavored at  $3.7\sigma$**

**Decoherence:**

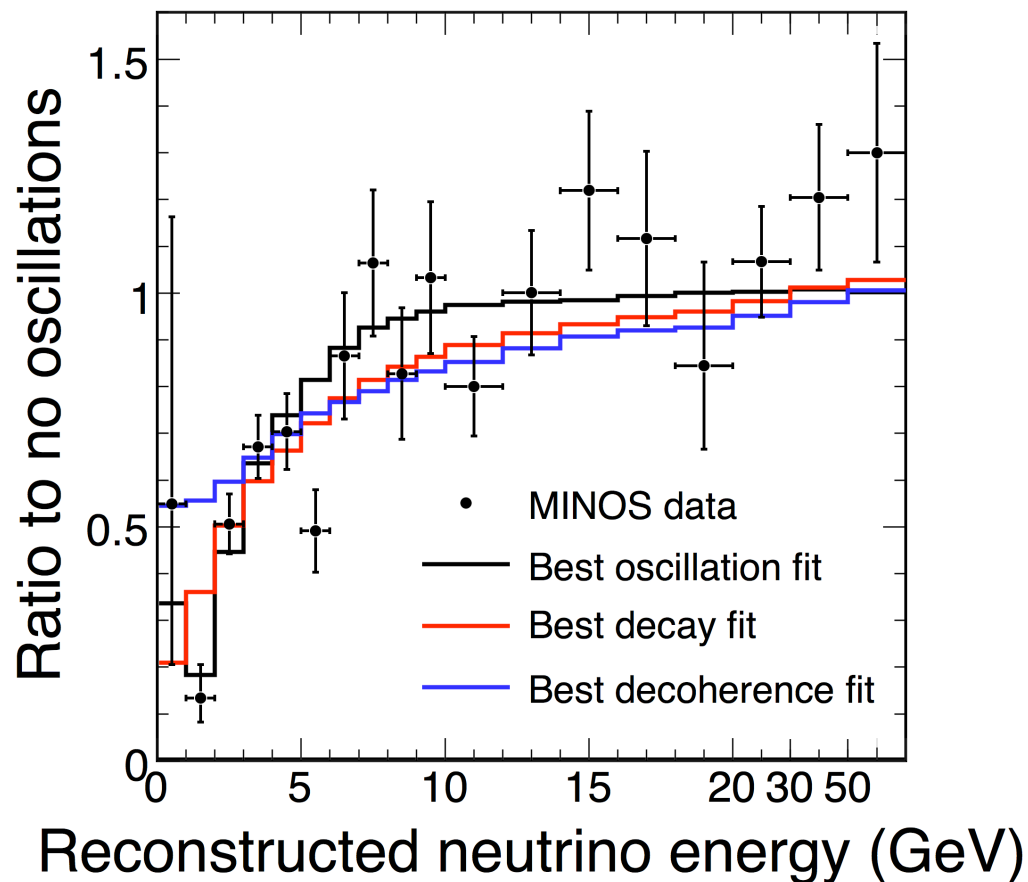
$$P_{\mu\mu} = 1 - \frac{\sin^2 2\theta}{2} \left( 1 - \exp\left(\frac{-\mu^2 L}{2E_\nu}\right) \right)$$

G.L. Fogli *et al.*, PRD67:093006 (2003)

$$\chi^2/\text{ndof} = 123/97$$

$$\Delta\chi^2 = 33$$

**disfavored at  $5.7\sigma$**



# Neutrino $\Delta m^2$ sensitivity evolution

